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Geodesic equations and their numerical solutions in geodetic and Cartesian coordinates on an oblate spheroid

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Abstract: The direct geodesic problem on an oblate spheroid is described as an initial value problem and is solved numerically using both geodetic and Cartesian coordinates. The geodesic equations are formulated by means of the theory of differential geometry. The initial value problem under consideration is reduced to a system of first-order ordinary differential equations, which is solved using a numerical method. The solution provides the coordinates and the azimuths at any point along the geodesic. The Clairaut constant is not used for the solution but it is computed, allowing to check the precision of the method. An extensive data set of geodesics is used, in order to evaluate the performance of the method in each coordinate system. The results for the direct geodesic problem are validated by comparison to Karney's method. We conclude that a complete, stable, precise, accurate and fast solution of the problem in Cartesian coordinates is accomplished.

Keywords: Clairaut's constant, Direct geodesic problem, Geometrical geodesy, Karney's method, Numerical method

1 Introduction

In geodesy, there are two traditional problems concerning geodesics on an oblate spheroid (ellipsoid of revolution): (i) the direct problem: given a point P_0 on an oblate spheroid, together with the azimuth α_0 and the geodesic distance s_{01} to a point P_1 , determine the point P_1 and the azimuth α_1 at this point, and (ii) the inverse problem:

given two points P_0 and P_1 on an oblate spheroid, determine the geodesic distance s_{01} between them and the azimuths α_0, α_1 at the end points.

These problems have a very long history and several different methods of solving them have been proposed by many researchers, as they are reported in a comprehensive list by Karney (2016) in GeographicLib. Also, some of the existing methods have been presented by Rapp (1993) and Deakin and Hunter (2010).

The methods of solving the above problems can be divided into two general categories: (i) using an auxiliary sphere, e.g., Bessel (1826), Rainsford (1955), Robbins (1962), Sodano (1965), Saito (1970), Vincenty (1975), Saito (1979), Bowring (1983), Karney (2013) and (ii) without using an auxiliary sphere, e.g., Kivioja (1971), Holmstrom (1976), Jank and Kivioja (1980), Thomas and Featherstone (2005), Panou (2013), Panou et al. (2013), Tseng (2014). The methods which use an auxiliary sphere are based on the classical work of Bessel (1826) and its modifications. On the other hand, the methods without using an auxiliary sphere are attacking the problems directly on an oblate spheroid. They are conceptually simpler and can be generalized in the case of a triaxial ellipsoid, as has been already presented by Holmstrom (1976) and Panou (2013).

The solution of the geodesic problems, with one of the above two methods, includes evaluating elliptic integrals or solving differential equations using: (i) approximate analytical methods, e.g., Vincenty (1975), Holmstrom (1976), Pittman (1986), Mai (2010), Karney (2013) or (ii) numerical methods, e.g., Saito (1979), Rollins (2010), Sjöberg (2012), Sjöberg and Shirazian (2012), Panou et al. (2013). The approximate analytical methods are usually based on the fact that an oblate spheroid deviates slightly from a sphere, so these methods essentially involve a truncated series expansion. On the other hand, numerical methods can be used for ellipsoids of arbitrary flattening. In addition, they do not require a change in the theoretical background with a modification of the computational environment. However, they may suffer from computational errors, which are reduced with the improvements in modern

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computational systems. Furthermore, since the numerical methods perform computations at many points along the geodesic, they can be used as a convenient and efficient approach to trace the geodesic. If tracing is not needed, an analytical method may be sufficient to give the results of the geodesic problems.

From the vast literature on geodesic problems, one notices that the evaluation of the performance of the geodesic algorithms is based on the aspects of stability, accuracy and computational speed. Of course, the execution time depends strongly on the programming environment and the computing platform used. Today, with the broad availability of high speed computers, the execution time does not longer play an essential role. With regard to stability, the algorithms should be stable in the domain of use, i.e. without limitations to the input data of the problem. Finally, they should provide results with high accuracy, depending on the demands of the application.

In this work, the direct geodesic problem and the corresponding equations, with independent variable s , are numerically solved as an initial value problem, directly on an oblate spheroid, using two coordinate systems: geodetic and Cartesian. We note that, Panou et al. (2013) have numerically solved the inverse geodesic problem and the corresponding equations, with independent variable λ , as a boundary value problem. The method presented here can be generalized in the case of a triaxial ellipsoid and can be used for arbitrary flattening. However, in order to evaluate the method, we limit the numerical applications to the case of the WGS84 oblate spheroid.

In general, there are several numerical methods for solving an initial value problem (see Hildebrand 1974). In this study, we will use a relatively simple and commonly applied fourth-order Runge-Kutta method (see Butcher 1987), which has been applied successfully to the geodesic boundary value problem (Panou 2013, Panou et al. 2013). In addition, we examine the performance (stability, precision, execution time) of the method in both coordinate systems.

Kivioja (1971) has solved the geodesic initial value problem by numerical integration of a system of two differential equations. Thomas and Featherstone (2005) improved Kivioja's method by altering the system of the two differential equations along the geodesic, in order to avoid the singularity when the geodesic passes through the vertices. Alternatively, in this study we propose the numerical integration of a system of four differential equations of a geodesic in geodetic coordinates, which is free from this singularity. Although there are more equations, from the solution we can determine the Clairaut constant, at any point along the geodesic, and thus we are able to check the

precision of the numerical integration. We should mention that other methods for the solution of the direct geodesic problem, such as Rollins (2010) and Sjöberg and Shirazian (2012), use the Clairaut constant as an a priori known quantity in the equations and they employ iterative techniques.

Inevitably, in curvilinear coordinates there are two poles (singularities) and hence all of the above methods, without modifications, can be ill-behaved. This does not happen in Cartesian coordinates and the algorithms based on them are insensitive to singularities, such as $\tan(\pi/2)$. Also, the Cartesian coordinate system can be easily related to other curvilinear systems, many formulas in this system are simpler, without numerical difficulties and computations do not demand the use of trigonometric functions, which can make computer processing slow, as pointed out by Felski (2011).

Using the calculus of variations, the geodesic equations in Cartesian coordinates were derived on a sphere by Fox (1987) and on a triaxial ellipsoid by Holmstrom (1976). Although the approximate analytical solution given by Holmstrom (1976) can be applied in the degenerate case of an oblate spheroid, it is of low precision, since the precision was not his primary consideration.

The method proposed in this work introduces a new approach to the geodesic initial value problem, both in terms of its theoretical basis and its numerical solution. Part of the solution constitutes the solution of the direct geodesic problem, where the position and the azimuth are determined only at the end point of the geodesic. In addition, the geodesic initial value problem directly provides the trace of the full path of the geodesic.

2 Geodesics in geodetic coordinates

The geodesic initial value problem, expressed in geodetic coordinates on an oblate spheroid, consists of determining a geodesic, parametrized by its arc length s , $\varphi = \varphi(s)$, $\lambda = \lambda(s)$, with azimuths $\alpha = \alpha(s)$ along it, which passes through a given point $P_0(\varphi(0), \lambda(0))$ in a known direction (given azimuth $\alpha_0 = \alpha(0)$) and has a certain length s_{01} .

2.1 Geodesic equations

We consider an oblate spheroid which, in geodetic coordinates (φ, λ) , is described parametrically by

$$x = N \cos \varphi \cos \lambda \quad (1)$$

$$y = N \cos \varphi \sin \lambda \quad (2)$$

$$z = N \left(1 - e^2\right) \sin \varphi \quad (3)$$

where φ ($-\pi/2 \leq \varphi \leq +\pi/2$) is the geodetic latitude, λ ($-\pi < \lambda \leq +\pi$) is the geodetic longitude and

$$N = \frac{a}{\left(1 - e^2 \sin^2 \varphi\right)^{1/2}} \quad (4)$$

is the radius of curvature in the prime vertical normal section, with a the major semiaxis and e the first eccentricity. Also, it holds that $e^2 = f(2 - f)$, where f is the flattening. In this parametrization, the elements of the first fundamental form are (Vermeer 2015)

$$E = \frac{a^2(1 - e^2)^2}{\left(1 - e^2 \sin^2 \varphi\right)^3} \quad (5)$$

$$F = 0 \quad (6)$$

$$G = \frac{a^2 \cos^2 \varphi}{1 - e^2 \sin^2 \varphi} \quad (7)$$

In Eq. (6), $F = 0$ indicates that the φ -curves (parallels) and λ -curves (meridians) are orthogonal. Also, $E \neq 0$ for all φ and $G = 0$ when $\varphi = \pm\pi/2$ (at the poles) (Panou et al. 2013). From Eqs. (3), we obtain the derivatives

$$E_\varphi = \frac{3e^2 a^2 (1 - e^2)^2 \sin(2\varphi)}{\left(1 - e^2 \sin^2 \varphi\right)^4} \quad (8)$$

$$F_\varphi = 0 \quad (9)$$

$$G_\varphi = -\frac{a^2(1 - e^2) \sin(2\varphi)}{\left(1 - e^2 \sin^2 \varphi\right)^2} \quad (10)$$

$$E_\lambda = F_\lambda = G_\lambda = 0 \quad (11)$$

Thus, the Christoffel symbols Γ_{jk}^i ($i, j, k = 1, 2$) become (Struik 1961, p. 107)

$$\Gamma_{11}^1 = \frac{E_\varphi}{2E} = \frac{3e^2 \sin(2\varphi)}{2\left(1 - e^2 \sin^2 \varphi\right)} \quad (12)$$

$$\Gamma_{22}^1 = -\frac{G_\varphi}{2E} = \frac{\left(1 - e^2 \sin^2 \varphi\right) \sin(2\varphi)}{2\left(1 - e^2\right)} \quad (13)$$

$$\Gamma_{12}^2 = \frac{G_\varphi}{2G} = -\frac{\left(1 - e^2\right) \tan \varphi}{1 - e^2 \sin^2 \varphi} \quad (14)$$

$$\Gamma_{12}^1 = \Gamma_{11}^2 = \Gamma_{22}^2 = 0 \quad (15)$$

Therefore, the geodesic equations, expressed in geodetic coordinates on an oblate spheroid, are given by (Struik 1961, p. 132)

$$\frac{d^2 \varphi}{ds^2} + \Gamma_{11}^1 \left(\frac{d\varphi}{ds}\right)^2 + \Gamma_{22}^1 \left(\frac{d\lambda}{ds}\right)^2 = 0 \quad (16)$$

$$\frac{d^2 \lambda}{ds^2} + 2\Gamma_{12}^2 \frac{d\varphi}{ds} \frac{d\lambda}{ds} = 0 \quad (17)$$

The initial conditions associated with these equations are

$$\varphi_0 = \varphi(0), \quad \left.\frac{d\varphi}{ds}\right|_0 = \frac{d\varphi}{ds}(0) \quad (18)$$

$$\lambda_0 = \lambda(0), \quad \left.\frac{d\lambda}{ds}\right|_0 = \frac{d\lambda}{ds}(0) \quad (19)$$

where the values of the derivatives at point $P_0(\varphi_0, \lambda_0)$ are given below. Hence, the direct geodesic problem is described as an initial value problem in geodetic coordinates on an oblate spheroid by Eqs. (16) and Eqs. (17).

2.2 Numerical solution

In order to solve the geodesic initial value problem by a numerical method, the system of two non-linear second-order ordinary differential equations (Eqs. (16) and (17)) are written equivalently as a system of four first-order differential equations:

$$\frac{d}{ds}(\varphi) = \frac{d\varphi}{ds} \quad (20)$$

$$\frac{d}{ds} \left(\frac{d\varphi}{ds}\right) = -\Gamma_{11}^1 \left(\frac{d\varphi}{ds}\right)^2 - \Gamma_{22}^1 \left(\frac{d\lambda}{ds}\right)^2 \quad (21)$$

$$\frac{d}{ds}(\lambda) = \frac{d\lambda}{ds} \quad (22)$$

$$\frac{d}{ds} \left(\frac{d\lambda}{ds}\right) = -2\Gamma_{12}^2 \frac{d\varphi}{ds} \frac{d\lambda}{ds} \quad (23)$$

This system can be integrated on the interval $[0, s]$ using a numerical method, such as Runge-Kutta (see Hildebrand 1974, Butcher 1987). The step size δs is given by $\delta s = s/n$, where n is the number of steps. As a rule, a greater number of steps leads to a greater precision but also to a greater execution time and vice versa. The effects of the number of steps and the performance of the method (stability, precision, execution time) are examined in detail in Section 4.

For the variables φ and λ , the initial conditions are φ_0 and λ_0 , respectively. For the required derivatives, we recall

the well-known relations of a differential element on an oblate spheroid, for any curve of length ds (Vermeer 2015)

$$\frac{d\varphi}{ds} = \frac{\cos \alpha}{M} \quad (24)$$

$$\frac{d\lambda}{ds} = \frac{\sin \alpha}{N \cos \varphi} \quad (25)$$

where

$$M = \frac{a(1-e^2)}{(1-e^2 \sin^2 \varphi)^{3/2}} \quad (26)$$

is the radius of curvature in the meridian normal section. Hence, the required values of the derivatives at point $P_0(\varphi_0, \lambda_0)$ are

$$\left. \frac{d\varphi}{ds} \right|_0 = \frac{\cos \alpha_0}{M(\varphi_0)} \quad (27)$$

$$\left. \frac{d\lambda}{ds} \right|_0 = \frac{\sin \alpha_0}{N(\varphi_0) \cos \varphi_0} \quad (28)$$

2.3 Azimuths and Clairaut's constant

Eqs. (24) and (25), the azimuth α at any point along the geodesic is computed by

$$\alpha = \arctan\left(\frac{V}{U}\right) = \operatorname{arccot}\left(\frac{U}{V}\right) \quad (29)$$

where

$$U = M \frac{d\varphi}{ds} \quad (30)$$

$$V = N \cos \varphi \frac{d\lambda}{ds} \quad (31)$$

Note that Eqs. (29) involves the variables φ , $d\varphi/ds$ and $d\lambda/ds$, which are obtained by the numerical integration.

The integration of Eq. (17) yields

$$\frac{d\lambda}{ds} G = C \quad (32)$$

where C is an arbitrary constant. We note that Eq. (32) involves only the variables φ and $d\lambda/ds$. Substituting Eq. (25) into Eq. (32) and using Eqs. (4) and (7), we obtain

$$N \cos \varphi \sin \alpha = C \quad (33)$$

which is the well-known Clairaut's equation in geodetic coordinates. Hence, Eq. (32) and Eq. (33) are mathematically equivalent and C is the Clairaut constant. Also, we can estimate, at any value of the independent variable s , the difference $\delta C = C - C_0$ between the computed value C and the known value C_0 at point P_0 , from the given φ_0 and α_0 , by means of Clairaut's equation (Eq. (33)). In this way, we can check the precision of the numerical integration, since the difference δC should be zero meters at any point along the geodesic.

3 Geodesics in Cartesian coordinates

In a similar manner, the geodesic initial value problem, expressed in Cartesian coordinates on an oblate spheroid, consists of determining a geodesic, parametrized by its arc length s , $x = x(s)$, $y = y(s)$, $z = z(s)$, with azimuths $\alpha = \alpha(s)$ along it, which passes through a given point $P_0(x(0), y(0), z(0))$ in a known direction (given azimuth $\alpha_0 = \alpha(0)$) and has a certain length s_{01} .

3.1 Geodesic equations

We consider an oblate spheroid which is described in Cartesian coordinates (x, y, z) by

$$S(x, y, z) = x^2 + y^2 + \frac{z^2}{1-e^2} - a^2 = 0 \quad (34)$$

It is well-known, from the theory of differential geometry, that the principal normal to the geodesic must coincide with the normal to the oblate spheroid (Struik 1961, Deakin and Hunter 2010), i.e.

$$\frac{d^2x/ds^2}{\partial S/\partial x} = \frac{d^2y/ds^2}{\partial S/\partial y} = \frac{d^2z/ds^2}{\partial S/\partial z} = -m \quad (35)$$

where m is a function of s . From these equations, together with Eq. (34), it is possible to determine $x(s)$, $y(s)$, $z(s)$ and $m(s)$. Using Eq. (34), Eqs. (35) become

$$\frac{1}{x} \frac{d^2x}{ds^2} = \frac{1}{y} \frac{d^2y}{ds^2} = \frac{1-e^2}{z} \frac{d^2z}{ds^2} = -2m \quad (36)$$

Differentiating Eq. (34), we have

$$x \frac{dx}{ds} + y \frac{dy}{ds} + \frac{z}{1-e^2} \frac{dz}{ds} = 0 \quad (37)$$

and a further differentiation yields

$$\begin{aligned} & x \frac{d^2x}{ds^2} + y \frac{d^2y}{ds^2} + \frac{z}{1-e^2} \frac{d^2z}{ds^2} \\ &= - \left[\left(\frac{dx}{ds} \right)^2 + \left(\frac{dy}{ds} \right)^2 + \frac{1}{1-e^2} \left(\frac{dz}{ds} \right)^2 \right] \end{aligned} \quad (38)$$

Hence, from Eqs. (36) and Eq. (38), we obtain

$$m = \frac{h}{2H} \quad (39)$$

where

$$H = x^2 + y^2 + \frac{z^2}{(1-e^2)^2} \quad (40)$$

$$h = \left(\frac{dx}{ds} \right)^2 + \left(\frac{dy}{ds} \right)^2 + \frac{1}{1-e^2} \left(\frac{dz}{ds} \right)^2 \quad (41)$$

Substituting Eq. (39) into Eqs. (36), we obtain the geodesic equations in Cartesian coordinates on an oblate spheroid

$$\frac{d^2x}{ds^2} + \frac{h}{H}x = 0 \tag{42}$$

$$\frac{d^2y}{ds^2} + \frac{h}{H}y = 0 \tag{43}$$

$$\frac{d^2z}{ds^2} + \frac{h}{H} \cdot \frac{z}{1 - e^2} = 0 \tag{44}$$

which are subject to the initial conditions

$$x_0 = x(0), \frac{dx}{ds} \Big|_0 = \frac{dx}{ds}(0) \tag{45}$$

$$y_0 = y(0), \frac{dy}{ds} \Big|_0 = \frac{dy}{ds}(0) \tag{46}$$

$$z_0 = z(0), \frac{dz}{ds} \Big|_0 = \frac{dz}{ds}(0) \tag{47}$$

where the values of the derivatives at point $P_0(x_0, y_0, z_0)$ are given below. Hence, the direct geodesic problem is described as an initial value problem in Cartesian coordinates on an oblate spheroid by Eqs. (42) to (47).

3.2 Numerical solution

In order to solve the above problem, the system of three non-linear second-order ordinary differential equations (Eqs. (42) to (44)) is rewritten as a system of six first-order differential equations:

$$\frac{d}{ds}(x) = \frac{dx}{ds} \tag{48}$$

$$\frac{d}{ds} \left(\frac{dx}{ds} \right) = -\frac{h}{H}x \tag{49}$$

$$\frac{d}{ds}(y) = \frac{dy}{ds} \tag{50}$$

$$\frac{d}{ds} \left(\frac{dy}{ds} \right) = -\frac{h}{H}y \tag{51}$$

$$\frac{d}{ds}(z) = \frac{dz}{ds} \tag{52}$$

$$\frac{d}{ds} \left(\frac{dz}{ds} \right) = -\frac{h}{H} \cdot \frac{z}{1 - e^2} \tag{53}$$

This system can be integrated on the interval $[0, s]$ by a numerical method. Again, the step size δs is given by $\delta s = s/n$, where n is the number of steps. For the variables x, y and z , the initial conditions are x_0, y_0 and z_0 , respectively.

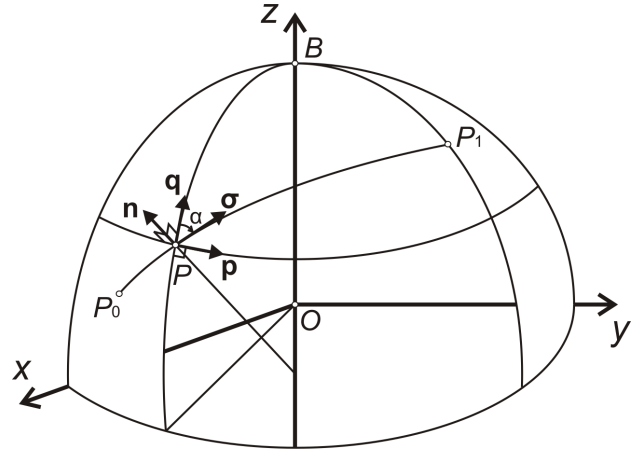


Figure 1: Unit vectors to a geodesic through a point P on an oblate spheroid: σ tangent to the geodesic, n normal to the oblate spheroid, p tangent to the parallel, q tangent to the meridian.

To obtain the required derivatives, we proceed to describe the unit vectors to a geodesic through a point $P(x, y, z)$ on an oblate spheroid (see Fig. 1).

Let σ be a unit vector tangent to an arbitrary geodesic through P . Then, we can express σ in terms of the unit vectors p, q and the azimuth α (Fig. 1):

$$\sigma = \left(\frac{dx}{ds}, \frac{dy}{ds}, \frac{dz}{ds} \right) = p \sin \alpha + q \cos \alpha \tag{54}$$

The unit vector normal to an oblate spheroid (using the gradient operator and Eqs. (34), (40)) can be expressed as (Fig. 1):

$$n = (n_1, n_2, n_3) = \left(\frac{x}{H^{1/2}}, \frac{y}{H^{1/2}}, \frac{z}{(1 - e^2) H^{1/2}} \right) \tag{55}$$

Furthermore, considering the plane of the meridian of the point P , which passes also through the pole B and the centre of the oblate spheroid O (Felski 2011), we obtain the unit vector p (Fig. 1):

$$p = (p_1, p_2, p_3) = \left(\frac{-y}{(x^2 + y^2)^{1/2}}, \frac{x}{(x^2 + y^2)^{1/2}}, 0 \right) \tag{56}$$

This vector has singularities at the poles, where we can simply set $p = (0, 1, 0)$. Otherwise, this vector can be expressed in terms of geodetic longitude λ with the help of Eqs. (1) and (2):

$$p = (p_1, p_2, p_3) = (-\sin \lambda, \cos \lambda, 0) \tag{57}$$

At the poles, we can now set $\lambda = \alpha$, i.e. $p = (-\sin \alpha, \cos \alpha, 0)$.

The unit vector q , tangent to the meridian, can now be determined as the cross product of unit vectors n and p

(Fig. 1):

$$\mathbf{q} = \mathbf{n} \times \mathbf{p} = (q_1, q_2, q_3) = (-n_3 p_2, n_3 p_1, n_1 p_2 - n_2 p_1) \quad (58)$$

Finally, substituting the vectors \mathbf{p} and \mathbf{q} into Eq. (54), we obtain the required values of the derivatives at point $P_0(x_0, y_0, z_0)$

$$\left. \frac{dx}{ds} \right|_0 = p_1(0) \sin \alpha_0 + q_1(0) \cos \alpha_0 \quad (59)$$

$$\left. \frac{dy}{ds} \right|_0 = p_2(0) \sin \alpha_0 + q_2(0) \cos \alpha_0 \quad (60)$$

$$\left. \frac{dz}{ds} \right|_0 = p_3(0) \sin \alpha_0 + q_3(0) \cos \alpha_0 \quad (61)$$

3.3 Azimuths and Clairaut's constant

Taking the scalar product of Eq. (54) successively with \mathbf{p} and \mathbf{q} and dividing the resulting equations, yields

$$\alpha = \arctan\left(\frac{R}{Q}\right) = \operatorname{arccot}\left(\frac{Q}{R}\right) \quad (62)$$

where

$$Q = \mathbf{q} \cdot \boldsymbol{\sigma} = q_1 \frac{dx}{ds} + q_2 \frac{dy}{ds} + q_3 \frac{dz}{ds} \quad (63)$$

$$R = \mathbf{p} \cdot \boldsymbol{\sigma} = p_1 \frac{dx}{ds} + p_2 \frac{dy}{ds} + p_3 \frac{dz}{ds} \quad (64)$$

Note that Eqs. (62) involves all the variables x , dx/ds , y , dy/ds , z and dz/ds , which are obtained by the numerical integration.

From the first equation of Eqs. (36), a second-order differential equation, we obtain a first integral

$$x \frac{dy}{ds} - y \frac{dx}{ds} = C \quad (65)$$

where C is an arbitrary constant. We note that Eq. (65) involves only the variables x , dx/ds , y and dy/ds . Now, from the scalar product of Eq. (54) with \mathbf{p} and using Eqs. (56), (65), we obtain

$$(x^2 + y^2)^{1/2} \sin \alpha = C \quad (66)$$

which is the well-known Clairaut's equation in Cartesian coordinates. Hence, Eq. (65) and Eq. (66) are mathematically equivalent and C is the Clairaut constant. Also, in a manner similar to the geodetic coordinates, we can estimate, at any value of s , the difference $\delta C = C - C_0$ between the computed value C and the known value C_0 at point P_0 , from the given x_0 , y_0 and α_0 , by means of Clairaut's equation (Eq. (66)). Furthermore, because the numerical

integration is performed in space, we can compute, at any value of s , the function S , given by Eq. (34). Therefore, we can double check the precision of the numerical integration, since the difference δC should be zero meters and the function S should be zero at any point along the geodesic on an oblate spheroid.

4 Numerical tests and comparisons

4.1 Test set

In order to evaluate the performance of the solution in each coordinate system (with respect to stability, precision and execution time) and to validate the method presented above, we used an extensive test set of geodesics, which is available in Karney (2010). This is a set of 500000 geodesics for the WGS84 ellipsoid of revolution, with $a = 6378137$ m and $f = 1/298.257223563$. The geodesics of the set are distributed into nine groups, as described in Table 1.

Table 1: Description of the geodesics in the test set.

Group	Identification number (ID)	Case
1	1 – 100000	randomly distributed
2	100001 – 150000	nearly antipodal
3	150001 – 200000	short distances
4	200001 – 250000	one end near a pole
5	250001 – 300000	both ends near opposite poles
6	300001 – 350000	nearly meridional
7	350001 – 400000	nearly equatorial
8	400001 – 450000	running between vertices ($\alpha_0 = \alpha_1 = 90^\circ$)
9	450001 – 500000	ending close to vertices

Each geodesic of the test set is defined by the data and the results for the direct geodesic problem, as described in Table 2.

The values of φ_1^K , λ_1^K and α_1^K were computed by Karney using high-precision direct geodesic calculations with the given φ_0 , λ_0 , α_0 and s_{01} . For simplicity and without loss of generality, φ_0 is chosen in $[0^\circ, 90^\circ]$, λ_0 is taken to be zero, α_0 is chosen in $[0^\circ, 180^\circ]$. Furthermore, φ_0 and α_0 are taken to be multiples of 10^{-12} deg and s_{01} is a multiple of 0.1 μm in $[0 \text{ m}, 20003931.4586254 \text{ m}]$. Also, the values for s_{01} for the geodesics running between vertices are truncated to a multiple of 0.1 pm and this is used to deter-

Table 2: Description of data and results for the direct geodesic problem in the test set.

Quantity	Symbol	Unit	Accuracy
latitude at point 0	φ_0	degrees	exact
longitude at point 0	λ_0	degrees	exact, always 0
azimuth at point 0	α_0	clockwise from north in degrees	exact
latitude at point 1	φ_1^K	degrees	accurate to 10^{-18} deg
longitude at point 1	λ_1^K	degrees	accurate to 10^{-18} deg
azimuth at point 1	α_1^K	degrees	accurate to 10^{-18} deg
geodesic distance from point 0 to point 1	s_{01}	meters	exact

mine point 1. Finally, these conditions result to having λ_1 in $[0^\circ, 180^\circ]$ and α_1 in $[0^\circ, 180^\circ]$ (Karney 2016).

For each geodesic in the test set, the systems of first-order differential equations (Eqs. (20) to (23) and Eqs. (48) to (53)) were integrated using the fourth-order Runge-Kutta numerical method (see Hildebrand 1974, Butcher 1987) using several different values of the number of steps n .

All algorithms were coded in FORTRAN95, were compiled by the open-source GNU FORTRAN compiler (at Level 2 optimization) and were executed on a personal computer running a 64-bit operating system. The main characteristics of the hardware were: Intel Core i5-2430M CPU (clocked at 2.4 GHz) and 6 GB of RAM.

For the computations, we used a 10-byte floating point arithmetic, which provides a precision of 18 decimal digits. However, for the conversion of the data of the full set (Table 2) from geodetic to Cartesian coordinates ($x_0, y_0 = 0, z_0, x_1^K, y_1^K, z_1^K$) using Eqs. (1) to (3), as well as for the computation of Clairaut's constant C_1^K at the end point using Eq. (33), a 16-byte arithmetic was employed, which provides a precision of 33 decimal digits. The same high-precision arithmetic was also used for several tests (e.g. Table 16).

Detailed results of solving the direct geodesic problem in both coordinate systems are presented in the following sections.

4.2 Solving the geodesics in geodetic coordinates

The performance of the proposed method, using geodetic coordinates, was evaluated through a sub-set of 338640 geodesics, having the composition shown in Table 3.

This sub-set was formed considering the constraints presented in Thomas and Featherstone (2005), in order to avoid the instabilities caused by the singularities in the geodetic coordinates. In particular, the geodesics of the

Table 3: Composition of the sub-set.

Group	Total number of geodesics
1	97997
2	49441
3	46699
4	0
5	0
6	0
7	50000
8	47294
9	47209

sub-set satisfy the following criteria: $\varphi_0 < 85^\circ$, $|\varphi_1| < 85^\circ$, $1^\circ < \alpha_0 < 179^\circ$ and $1^\circ < \alpha_1 < 179^\circ$.

The direct geodesic problem in geodetic coordinates is solved using the input data $\varphi_0, \lambda_0, \alpha_0$ and s_{01} . The results (φ_1, λ_1) at the end point are converted to Cartesian coordinates (x_1, y_1, z_1) because this transformation is simple and numerically stable. In addition, this permits a direct comparison with the corresponding results using Cartesian coordinates.

At any point along the geodesic, the difference $\delta C = C - C_0$ is computed and its maximum value $\max|\delta C|$ is recorded. We also record the $\max(\max|\delta C|)$ of the whole set.

At the end point of each geodesic, we compute the difference $\delta r_1 = \left[(x_1 - x_1^K)^2 + (y_1 - y_1^K)^2 + (z_1 - z_1^K)^2 \right]^{1/2}$ and we record the $\max(\delta r_1)$ of the whole set. Similarly, we compute the differences $\delta \alpha_1 = \alpha_1 - \alpha_1^K$ and $\delta C_1 = C_1 - C_1^K$ and we record the $\max|\delta \alpha_1|$ and the $\max|\delta C_1|$. Finally, we record the execution time t , which was needed using the specified computing system.

We note that we do not compute the azimuth at the intermediate points of the geodesic (Eqs. (29)), since we have no other similar data to compare with.

All of the above data, along with the ID of the relevant geodesic, are presented in Table 4, for several values of the number of steps of the integration.

4.3 Solving the geodesics in Cartesian coordinates

In a similar manner, Table 5 presents the corresponding results of the performance of the proposed method in Cartesian coordinates, using the same sub-set of 338640 geodesics and using several values for the number of steps in the numerical integration. The input data are now x_0 , y_0 , z_0 , α_0 and s_{01} .

Comparing the results presented in Table 4 and Table 5, it is remarkable that, using Cartesian coordinates, we achieve similar or better levels of accuracy with a much smaller number of integration steps than using geodetic coordinates. Therefore, the execution time in Cartesian coordinates is reduced by a factor about 60, for similar accuracy levels.

Table 6 describes the results obtained solving the full set of 500000 geodesics in Cartesian coordinates. We note that, at any point along each geodesic, we compute the absolute value of the function S (Eq. (34)) and we record the $\max|S|$. Thus, in addition to the data presented in the previous tables, we also give the value of $\max(\max|S|)$ for the whole set.

From Table 6 we conclude that our results are in agreement with the results of Karney's method, within a few nanometers in the end position and a few microseconds in the end azimuth.

With regard to the execution time, these impressive results are obtained at an average rate of 0.3 μ s per integration step, which corresponds to about 6 ms for a geodesic using 20000 points. However, for most practical applications, a smaller number of steps is quite adequate.

In order to study in detail the differences in the results at the end points, we subsequently present, in Tables 7 to 15, the main results, separately for each group.

We remark that the execution times for groups 2 to 9 (all have 50000 geodesics) were almost identical, so we present execution times only for group 1 (100000 geodesics) and group 2.

In Table 9 one may notice that a very small number of integration steps is sufficient to provide accurate results in the case of very short geodesics. In this case, an increase in the number of steps leads to worse results, which are attributed to the effects of round-off errors.

Since the results for groups 5, 8 and 9 indicate a lower accuracy of the value $\max|\delta\alpha_1|$, we use a larger number of

steps (50000) but the improvement is small, especially for group 5.

In order to examine further the cause of this behavior, we solved a particularly ill-behaved geodesic (ID 294750) using high-precision arithmetic (33 digits) and greater numbers of integration steps. The results, which are presented in Table 16, show a full agreement with those of Karney's method.

5 Concluding remarks

A numerical solution of the geodesic initial value problem in geodetic and Cartesian coordinates on an oblate spheroid has been presented. The real power of the proposed method is that it is universal, i.e. can be used for arbitrary flattening and can be generalized in the case of a triaxial ellipsoid. Also, by setting $e = 0$ in the formulations, the geodesic initial value problem and its numerical solution in geodetic and Cartesian coordinates on a sphere is obtained as a degenerate case.

Comparing the results of solving the geodesic initial value problem in the two coordinate systems, we conclude that only the solution in Cartesian coordinates is complete, i.e. it works in the entire range of input data, it is stable, precise, accurate and fast, so it is recommended for use, especially when a tracing of the geodesic line is required. The precision of the method depends on the number of the significant digits of the computer system being used. However, current computer systems of everyday use are adequate to achieve excellent results in a very short time.

Furthermore, employing higher precision arithmetic (larger number of decimal digits), the results obtained, using the proposed method in Cartesian coordinates, are directly comparable with the results of Karney's method and they constitute an independent validation and verification of the test set of Karney (2010).

We also investigate ways to improve the performance of the proposed method, with regard to the precision attained in relation to the required execution time. We are experimenting with different orders of the numerical integration method, a more detailed examination of the dependence on the number of integration steps, as well as with using a variable step algorithm. We are also working on the generalization of this method and its application to a triaxial ellipsoid.

For the sake of a complete theory, knowledge of an analytical solution of the direct geodesic problem, in Cartesian coordinates, is of interest. Although today, with Global Navigation Satellite System (GNSS) technologies,

Table 4: Performance of the method on the subset of 338640 geodesics using geodetic coordinates.

n	$\max(\max \delta C)$ (m)	ID	$\max(\delta r_1)$ (m)	ID	$\max \delta\alpha_1 $ (arcsec)	ID	$\max \delta C_1 $ (m)	ID	t (s)
1000	42	110733	$1.5 \cdot 10^4$	145577	81	145577	$3.2 \cdot 10^2$	110733	302
2000	2.7	110733	$3.0 \cdot 10^2$	145577	1.7	145577	$1.5 \cdot 10^1$	110733	618
5000	$3.4 \cdot 10^{-2}$	110733	4.4	145577	$2.3 \cdot 10^{-2}$	145577	$3.7 \cdot 10^{-1}$	110733	1551
10000	$1.1 \cdot 10^{-3}$	110733	$1.7 \cdot 10^{-1}$	140863	$7.4 \cdot 10^{-4}$	145577	$2.2 \cdot 10^{-2}$	110733	3051
20000	$3.5 \cdot 10^{-5}$	110733	$9.1 \cdot 10^{-3}$	140863	$2.2 \cdot 10^{-5}$	145577	$1.4 \cdot 10^{-3}$	110733	5659
50000	$3.6 \cdot 10^{-7}$	110733	$2.2 \cdot 10^{-4}$	140863	$2.2 \cdot 10^{-7}$	110733	$3.5 \cdot 10^{-5}$	110733	14768
100000	$2.8 \cdot 10^{-8}$	170970	$1.4 \cdot 10^{-5}$	140863	$1.3 \cdot 10^{-8}$	110733	$2.2 \cdot 10^{-6}$	110733	28792
150000	$4.7 \cdot 10^{-8}$	154108	$2.7 \cdot 10^{-6}$	140863	$5.6 \cdot 10^{-9}$	154108	$4.3 \cdot 10^{-7}$	110733	43621

Table 5: Performance of the method on the subset of 338640 geodesics using Cartesian coordinates.

n	$\max(\max \delta C)$ (m)	ID	$\max(\delta r_1)$ (m)	ID	$\max \delta\alpha_1 $ (arcsec)	ID	$\max \delta C_1 $ (m)	ID	t (s)
500	$2.7 \cdot 10^{-6}$	490402	$2.6 \cdot 10^{-4}$	126166	$9.6 \cdot 10^{-5}$	424861	$2.7 \cdot 10^{-6}$	490402	75
1000	$8.3 \cdot 10^{-8}$	434510	$1.6 \cdot 10^{-5}$	126166	$6.0 \cdot 10^{-6}$	424861	$8.3 \cdot 10^{-8}$	434510	123
2000	$2.6 \cdot 10^{-9}$	462032	$1.0 \cdot 10^{-6}$	126166	$3.7 \cdot 10^{-7}$	424861	$2.6 \cdot 10^{-9}$	462032	226
5000	$9.1 \cdot 10^{-10}$	159630	$2.6 \cdot 10^{-8}$	418118	$9.6 \cdot 10^{-9}$	499415	$9.1 \cdot 10^{-10}$	159630	478
10000	$1.4 \cdot 10^{-9}$	159630	$2.2 \cdot 10^{-9}$	159630	$6.3 \cdot 10^{-10}$	443838	$1.4 \cdot 10^{-9}$	159630	928

Table 6: Performance of the method on the full set of 500000 geodesics.

n	$\max(\max \delta C)$ (m)	ID	$\max(\max S)$	ID	$\max(\delta r_1)$ (m)	ID	$\max \delta\alpha_1 $ (arcsec)	ID	$\max \delta C_1 $ (m)	ID	t (s)
100	$8.3 \cdot 10^{-3}$	413757	$1.4 \cdot 10^{-9}$	130433	$1.6 \cdot 10^{-1}$	117312	$4.3 \cdot 10^3$	293277	$8.3 \cdot 10^{-3}$	413757	83
200	$2.6 \cdot 10^{-4}$	407322	$4.3 \cdot 10^{-11}$	130433	$1.0 \cdot 10^{-2}$	117312	$2.7 \cdot 10^2$	293277	$2.6 \cdot 10^{-4}$	407322	106
500	$2.7 \cdot 10^{-6}$	490402	$4.5 \cdot 10^{-13}$	261766	$2.6 \cdot 10^{-4}$	117312	6.9	293277	$2.7 \cdot 10^{-6}$	490402	132
1000	$8.3 \cdot 10^{-8}$	434510	$1.0 \cdot 10^{-14}$	252658	$1.6 \cdot 10^{-5}$	289531	$4.3 \cdot 10^{-1}$	293277	$8.3 \cdot 10^{-8}$	434510	217
2000	$2.6 \cdot 10^{-9}$	462032	$< 10^{-14}$	296969	$1.0 \cdot 10^{-6}$	286391	$2.7 \cdot 10^{-2}$	293277	$2.6 \cdot 10^{-9}$	462032	358
5000	$9.1 \cdot 10^{-10}$	159630	$< 10^{-14}$	486112	$2.6 \cdot 10^{-8}$	299239	$7.0 \cdot 10^{-4}$	292563	$9.1 \cdot 10^{-10}$	159630	789
10000	$1.4 \cdot 10^{-9}$	159630	$< 10^{-14}$	159630	$2.2 \cdot 10^{-9}$	159630	$4.6 \cdot 10^{-5}$	292563	$1.4 \cdot 10^{-9}$	159630	1502
20000	$3.9 \cdot 10^{-9}$	193167	$< 10^{-14}$	193167	$4.2 \cdot 10^{-9}$	193167	$7.7 \cdot 10^{-6}$	293277	$3.9 \cdot 10^{-9}$	193167	2971

Table 7: Performance of the method for Group 1.

n	$\max(\delta r_1)$ (m)	$\max \delta\alpha_1 $ (arcsec)	$\max \delta C_1 $ (m)	t (s)
100	$1.6 \cdot 10^{-1}$	$2.4 \cdot 10^{-1}$	$7.7 \cdot 10^{-3}$	9
200	$1.0 \cdot 10^{-2}$	$1.5 \cdot 10^{-2}$	$2.4 \cdot 10^{-4}$	12
500	$2.6 \cdot 10^{-4}$	$3.8 \cdot 10^{-4}$	$2.5 \cdot 10^{-6}$	21
1000	$1.6 \cdot 10^{-5}$	$2.4 \cdot 10^{-5}$	$7.7 \cdot 10^{-8}$	34
2000	$1.0 \cdot 10^{-6}$	$1.5 \cdot 10^{-6}$	$2.4 \cdot 10^{-9}$	63
5000	$2.6 \cdot 10^{-8}$	$3.8 \cdot 10^{-8}$	$5.1 \cdot 10^{-11}$	155
10000	$1.7 \cdot 10^{-9}$	$2.4 \cdot 10^{-9}$	$6.3 \cdot 10^{-11}$	306
20000	$3.2 \cdot 10^{-10}$	$2.2 \cdot 10^{-10}$	$9.5 \cdot 10^{-11}$	553

Table 8: Performance of the method for Group 2.

n	$\max(\delta r_1)$ (m)	$\max \delta\alpha_1 $ (arcsec)	$\max \delta C_1 $ (m)	t (s)
100	$1.6 \cdot 10^{-1}$	$1.3 \cdot 10^{-1}$	$8.3 \cdot 10^{-3}$	5
200	$1.0 \cdot 10^{-2}$	$8.2 \cdot 10^{-3}$	$2.6 \cdot 10^{-4}$	6
500	$2.6 \cdot 10^{-4}$	$2.1 \cdot 10^{-4}$	$2.7 \cdot 10^{-6}$	10
1000	$1.6 \cdot 10^{-5}$	$1.3 \cdot 10^{-5}$	$8.3 \cdot 10^{-8}$	17
2000	$1.0 \cdot 10^{-6}$	$8.2 \cdot 10^{-7}$	$2.6 \cdot 10^{-9}$	31
5000	$2.6 \cdot 10^{-8}$	$2.1 \cdot 10^{-8}$	$6.9 \cdot 10^{-11}$	77
10000	$1.8 \cdot 10^{-9}$	$1.3 \cdot 10^{-9}$	$6.0 \cdot 10^{-11}$	150
20000	$3.9 \cdot 10^{-10}$	$8.9 \cdot 10^{-11}$	$8.5 \cdot 10^{-11}$	277

Table 9: Performance of the method for Group 3.

n	$\max(\delta r_1)$ (m)	$\max \delta\alpha_1 $ (arcsec)	$\max \delta C_1 $ (m)
100	$1.5 \cdot 10^{-11}$	$3.1 \cdot 10^{-13}$	$1.6 \cdot 10^{-11}$
200	$4.1 \cdot 10^{-11}$	$3.6 \cdot 10^{-13}$	$2.8 \cdot 10^{-11}$
500	$9.7 \cdot 10^{-11}$	$1.2 \cdot 10^{-12}$	$6.5 \cdot 10^{-11}$
1000	$1.6 \cdot 10^{-10}$	$1.2 \cdot 10^{-12}$	$1.3 \cdot 10^{-10}$
2000	$4.6 \cdot 10^{-10}$	$3.9 \cdot 10^{-12}$	$4.5 \cdot 10^{-10}$
5000	$1.1 \cdot 10^{-9}$	$5.7 \cdot 10^{-12}$	$9.1 \cdot 10^{-10}$
10000	$2.2 \cdot 10^{-9}$	$5.5 \cdot 10^{-12}$	$1.4 \cdot 10^{-9}$
20000	$4.2 \cdot 10^{-9}$	$5.3 \cdot 10^{-11}$	$3.9 \cdot 10^{-9}$

Table 10: Performance of the method for Group 4.

n	$\max(\delta r_1)$ (m)	$\max \delta\alpha_1 $ (arcsec)	$\max \delta C_1 $ (m)
100	$1.6 \cdot 10^{-1}$	$2.2 \cdot 10^{-3}$	$1.2 \cdot 10^{-6}$
200	$9.9 \cdot 10^{-3}$	$1.4 \cdot 10^{-4}$	$3.6 \cdot 10^{-8}$
500	$2.5 \cdot 10^{-4}$	$3.6 \cdot 10^{-6}$	$3.7 \cdot 10^{-10}$
1000	$1.6 \cdot 10^{-5}$	$2.2 \cdot 10^{-7}$	$1.2 \cdot 10^{-11}$
2000	$9.9 \cdot 10^{-7}$	$1.4 \cdot 10^{-8}$	$1.0 \cdot 10^{-11}$
5000	$2.5 \cdot 10^{-8}$	$3.6 \cdot 10^{-10}$	$1.5 \cdot 10^{-11}$
10000	$1.6 \cdot 10^{-9}$	$2.2 \cdot 10^{-11}$	$2.3 \cdot 10^{-11}$
20000	$2.6 \cdot 10^{-10}$	$1.8 \cdot 10^{-11}$	$2.9 \cdot 10^{-11}$

Table 11: Performance of the method for Group 5.

n	$\max(\delta r_1)$ (m)	$\max \delta\alpha_1 $ (arcsec)	$\max \delta C_1 $ (m)
100	$1.6 \cdot 10^{-1}$	$4.3 \cdot 10^3$	$1.5 \cdot 10^{-6}$
200	$1.0 \cdot 10^{-2}$	$2.7 \cdot 10^2$	$4.6 \cdot 10^{-8}$
500	$2.6 \cdot 10^{-4}$	6.9	$4.7 \cdot 10^{-10}$
1000	$1.6 \cdot 10^{-5}$	$4.3 \cdot 10^{-1}$	$1.6 \cdot 10^{-11}$
2000	$1.0 \cdot 10^{-6}$	$2.7 \cdot 10^{-2}$	$9.7 \cdot 10^{-12}$
5000	$2.6 \cdot 10^{-8}$	$7.0 \cdot 10^{-4}$	$1.5 \cdot 10^{-11}$
10000	$1.8 \cdot 10^{-9}$	$4.6 \cdot 10^{-5}$	$1.9 \cdot 10^{-11}$
20000	$3.5 \cdot 10^{-10}$	$7.7 \cdot 10^{-6}$	$3.0 \cdot 10^{-11}$
50000	$4.9 \cdot 10^{-10}$	$3.6 \cdot 10^{-6}$	$4.7 \cdot 10^{-11}$

Table 12: Performance of the method for Group 6.

n	$\max(\delta r_1)$ (m)	$\max \delta\alpha_1 $ (arcsec)	$\max \delta C_1 $ (m)
100	$1.6 \cdot 10^{-1}$	17	$6.9 \cdot 10^{-7}$
200	$1.0 \cdot 10^{-2}$	1.1	$2.2 \cdot 10^{-8}$
500	$2.6 \cdot 10^{-4}$	$2.7 \cdot 10^{-2}$	$2.2 \cdot 10^{-10}$
1000	$1.6 \cdot 10^{-5}$	$1.7 \cdot 10^{-3}$	$6.9 \cdot 10^{-12}$
2000	$1.0 \cdot 10^{-6}$	$1.1 \cdot 10^{-4}$	$2.2 \cdot 10^{-13}$
5000	$2.6 \cdot 10^{-8}$	$2.7 \cdot 10^{-6}$	$< 10^{-14}$
10000	$1.7 \cdot 10^{-9}$	$1.9 \cdot 10^{-7}$	$< 10^{-14}$
20000	$3.6 \cdot 10^{-10}$	$2.5 \cdot 10^{-9}$	$< 10^{-14}$

Table 13: Performance of the method for Group 7.

n	$\max(\delta r_1)$ (m)	$\max \delta\alpha_1 $ (arcsec)	$\max \delta C_1 $ (m)
100	$1.6 \cdot 10^{-1}$	$4.5 \cdot 10^{-7}$	$8.3 \cdot 10^{-3}$
200	$1.0 \cdot 10^{-2}$	$2.8 \cdot 10^{-8}$	$2.6 \cdot 10^{-4}$
500	$2.6 \cdot 10^{-4}$	$7.2 \cdot 10^{-10}$	$2.7 \cdot 10^{-6}$
1000	$1.6 \cdot 10^{-5}$	$4.5 \cdot 10^{-11}$	$8.3 \cdot 10^{-8}$
2000	$1.0 \cdot 10^{-6}$	$2.8 \cdot 10^{-12}$	$2.6 \cdot 10^{-9}$
5000	$2.6 \cdot 10^{-8}$	$1.1 \cdot 10^{-13}$	$6.4 \cdot 10^{-11}$
10000	$1.7 \cdot 10^{-9}$	$4.0 \cdot 10^{-14}$	$7.1 \cdot 10^{-11}$
20000	$3.3 \cdot 10^{-10}$	$4.0 \cdot 10^{-14}$	$1.0 \cdot 10^{-10}$

Table 14: Performance of the method for Group 8.

n	$\max(\delta r_1)$ (m)	$\max \delta\alpha_1 $ (arcsec)	$\max \delta C_1 $ (m)
100	$1.6 \cdot 10^{-1}$	36	$8.3 \cdot 10^{-3}$
200	$1.0 \cdot 10^{-2}$	2.3	$2.6 \cdot 10^{-4}$
500	$2.6 \cdot 10^{-4}$	$5.8 \cdot 10^{-2}$	$2.7 \cdot 10^{-6}$
1000	$1.6 \cdot 10^{-5}$	$3.6 \cdot 10^{-3}$	$8.3 \cdot 10^{-8}$
2000	$1.0 \cdot 10^{-6}$	$2.3 \cdot 10^{-4}$	$2.6 \cdot 10^{-9}$
5000	$2.6 \cdot 10^{-8}$	$5.8 \cdot 10^{-6}$	$7.4 \cdot 10^{-11}$
10000	$1.8 \cdot 10^{-9}$	$3.6 \cdot 10^{-7}$	$7.3 \cdot 10^{-11}$
20000	$3.6 \cdot 10^{-10}$	$3.7 \cdot 10^{-8}$	$9.4 \cdot 10^{-11}$
50000	$5.1 \cdot 10^{-10}$	$2.0 \cdot 10^{-8}$	$1.7 \cdot 10^{-10}$

Table 15: Performance of the method for Group 9.

n	$\max(\delta r_1)$ (m)	$\max \delta\alpha_1 $ (arcsec)	$\max \delta C_1 $ (m)
100	$1.6 \cdot 10^{-1}$	65	$8.3 \cdot 10^{-3}$
200	$1.0 \cdot 10^{-2}$	4.1	$2.6 \cdot 10^{-4}$
500	$2.6 \cdot 10^{-4}$	$1.0 \cdot 10^{-1}$	$2.7 \cdot 10^{-6}$
1000	$1.6 \cdot 10^{-5}$	$6.5 \cdot 10^{-3}$	$8.3 \cdot 10^{-8}$
2000	$1.0 \cdot 10^{-6}$	$4.1 \cdot 10^{-4}$	$2.6 \cdot 10^{-9}$
5000	$2.6 \cdot 10^{-8}$	$1.0 \cdot 10^{-5}$	$6.9 \cdot 10^{-11}$
10000	$1.8 \cdot 10^{-9}$	$6.5 \cdot 10^{-7}$	$6.3 \cdot 10^{-11}$
20000	$4.1 \cdot 10^{-10}$	$6.7 \cdot 10^{-8}$	$9.0 \cdot 10^{-11}$
50000	$4.6 \cdot 10^{-10}$	$2.4 \cdot 10^{-9}$	$1.5 \cdot 10^{-10}$

Table 16: Comparisons with Karney’s data for the geodesic 294750 using 33-digit arithmetic.

n	$\max(\delta r_1)$ (m)	$\max \delta\alpha_1 $ (arcsec)	t (s)
50000	$2.7 \cdot 10^{-12}$	$4.1 \cdot 10^{-8}$	0.6
100000	$2.1 \cdot 10^{-13}$	$2.6 \cdot 10^{-9}$	1.2
200000	$6.0 \cdot 10^{-14}$	$1.6 \cdot 10^{-10}$	2.5
500000	$5.0 \cdot 10^{-14}$	$4.1 \cdot 10^{-12}$	6.2
1000000	$<10^{-14}$	$2.5 \cdot 10^{-13}$	12.2
2000000	$<10^{-14}$	$1.0 \cdot 10^{-14}$	25.7

the inverse geodesic problem is more realistic than the direct geodesic problem, the proposed algorithm will be used iteratively for the solution of the inverse problem, as has already been suggested by Jank and Kivioja (1980) and Vermeer (2015).

Finally, we plan to apply this concept, i.e. to solve the problem in space rather than on a surface, to other curves of geodetic importance, such as the normal section curve, the curve of alignment, the great elliptic arc and the loxodrome, as well as in other suitable geodetic problems.

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