Theory of the Viscomagnetic Heat Flux

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The heat transport in the presence of a viscous flow and a magnetic field is studied for a rarefied polyatomic gas enclosed between infinite parallel plates. The total "viscomagnetic heat flux" is the sum of a bulk and a surface heat flux. According to moment equations of the Waldmann-Snider kinetic equation, the bulk heat flux is due to gradients of pressure (i. e. second order derivatives of flow velocity) and of tensor polarization. For the tensor polarization a second order differential equation is solved with boundary conditions obtained by Waldmann's thermodynamic method. This method also supplies an ansatz for the surface heat flux. In a thermally isolated box the viscomagnetic heat flux gives rise to a temperature gradient in y -direction if the magnetic field is in z -direction and the gas is flowing in x -direction. The bulk contributions to this transverse temperature gradient have already been given by Levi, McCourt and Beenakker. Now, the surface contributions and the total effect are studied quantitatively for some gases. In particular, it is discussed how a new surface parameter could be determined from future experimental data.

1. Introduction

Since Maxwell's work¹ on transport phenomena in rarefied gases it is known that modifications on the heat flux and on the friction pressure tensor occur at low pressures: Second order derivatives of flow velocity contribute to the heat flux², and likewise second gradients of temperature are present in the friction pressure tensor^{3,4}. The corresponding expressions for the "viscous heat flux" and the "thermal stresses" for a rarefied polyatomic gas in a magnetic field have been derived by Levi et al.^{5,6}. For a polyatomic gas in a magnetic field, the transport processes are no longer isotropic, the second rank thermal conductivity tensor and the fourth rank viscosity tensor replace the scalar transport coefficients: Senftleben-Beenakker effect⁷. Similarly, in rarefied gases the effects due to second order gradients of temperature and velocity are influenced by a magnetic field⁶. In particular, the magnetic field dependent thermal stresses become manifest^{5,6} in the Scott effect⁸ (thermomagnetic torque). As a reciprocal phenomenon, Levi et al.⁶ predicted the existence of a viscomagnetic heat flux: In a rarefied polyatomic gas, flowing in x -direction, a heat flux in y -direction should occur, if a magnetic field in z -direction is present. In a thermally isolated rectangular capillary, a transverse temperature gradient builds up. Recently, the existence of this effect has been demonstrated by Eggermont et al.⁹.

Besides the bulk effects, discussed so far, the boundary layer phenomena play an equally important role in a rarefied gas. So, for the thermomagnetic torque, slip effects (thermal⁵ and thermomagnetic¹⁰) have to be taken into account besides the bulk thermal stresses. This is in accordance with the existence of the thermomagnetic pressure difference^{11,12}, a phenomenon which is entirely determined by slip effects^{10,13}.

The topic of this paper is the calculation of the boundary layer contributions to the viscomagnetic heat flux. Two types of such contributions are discussed. First, it is noted that the total heat transported through the slit formed by two infinite parallel plates consists of the bulk heat flux $\mathbf{q}$ and of the "surface heat flux" $\mathbf{Q}$. The surface heat flux¹⁴⁻¹⁶ accounts for the difference between the actual value of the heat flux at the wall and the value of $\mathbf{q}$ extrapolated to the wall. Secondly, the bulk heat flux $\mathbf{q}$ itself is influenced by boundary effects. In this connection, the behaviour near the wall of the tensor polarization and the Kagan polarization (which are used for the description of the field effects) is of crucial importance. The treatment is based¹⁷

- i) on moment equations derived from the Waldmann-Snider kinetic equation^{18,19}, and
- ii) on boundary conditions obtained from the interfacial entropy production by Waldmann's thermodynamical method¹⁴.

In particular, a differential equation and a boundary condition for the tensor polarization, and an ansatz for the surface heat flux are available.

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The transverse temperature gradient calculated in this way consists of the bulk contribution already given by Levi et al.⁶ and of a "surface contribution". For the gas N_2 as a representative example, the importance of the surface contribution is studied quantitatively. To this end, for two of the phenomenological surface parameters the values are used which were obtained from the analysis of thermomagnetic pressure difference data¹². The thermal slip number is taken¹⁶ from thermal transpiration data. Furthermore, a new surface parameter occurs which characterizes the coupling of the surface heat flux with the boundary value of the Kagan polarization. It is shown how the magnitude of this parameter influences the total effect. Special attention is given to the value and position of the maximum of the transverse temperature gradient as a function of the magnetic field strength for different orientations of the field.

2. Statement of the Problem

In a rarefied polyatomic gas which flows through a flat rectangular capillary placed in a magnetic field, a transverse temperature gradient perpendicular to the flow direction occurs. This was predicted by Levi et al.⁶ and recently has experimentally been verified by Eggermont et al.⁹. For simplicity, theoretical calculations will be made for the flow of a gas between infinite parallel plates (of distance $2d$) located at $z = \pm d$, with a pressure gradient

$$\nabla p^{(0)} = (\partial p / \partial x)_0 \mathbf{e}_x$$

applied (for zero field) in x -direction.

The total heat transported parallel to the plates is the sum of a bulk and a surface term,

$$\mathbf{J}_q = \int_{-d}^{+d} dz \mathbf{q}^{\text{tan}} + \mathbf{Q}(d) + \mathbf{Q}(-d), \quad (1)$$

with $\mathbf{q}^{\text{tan}} \equiv \mathbf{q} - \mathbf{e}_z \mathbf{e}_z \cdot \mathbf{q}$. Here, $\mathbf{q}$ is the heat flux in the bulk gas and $\mathbf{Q}(\pm d)$ is the surface heat flux at the plates $z = \pm d$. The surface heat flux¹⁴ makes up for the difference between the true boundary value of the heat flux and the value of $\mathbf{q}$ obtained by extrapolation to the wall. From a thermodynamical point of view, the surface heat flux is linked with the thermal slip, i.e. with phenomena like thermal transpiration and thermal force¹⁶.

The bulk heat flux $\mathbf{q}$ is a sum of two terms,

$$\mathbf{q} = -\lambda \cdot \nabla T + \mathbf{q}_{\text{VM}}, \quad (2)$$

the ordinary contribution due to a temperature gradient, and the viscomagnetic bulk heat flux $\mathbf{q}_{\text{VM}}$ which arises from a pressure gradient, i.e. from second order derivatives of velocity. The second rank thermal conductivity tensor λ depends on the strength and orientation of the magnetic field, with the field influence (Senftleben-Beenakker effect⁷) being a small effect of the order of 1%. For zero field, the temperature is constant in plane Poiseuille flow. Hence, due to the smallness of the field dependent temperature gradient, the first term in Eq. (2) can in good approximation be replaced by its isotropic value $-\lambda \nabla T$:

$$\mathbf{q} \approx -\lambda \nabla T + \mathbf{q}_{\text{VM}}. \quad (3)$$

Under the assumption that ∇T depends on z only, both $\partial T / \partial x$ and $\partial T / \partial y$ are constant, so Eqs. (1) and (3) give

$$\mathbf{J}_q \approx 2d [-\lambda (\nabla T)^{\text{tan}} + \overline{\mathbf{q}_{\text{VM}}^{\text{tan}}}] + \mathbf{Q}(d) + \mathbf{Q}(-d), \quad (4)$$

where

$$\overline{\mathbf{q}_{\text{VM}}^{\text{tan}}} = \frac{1}{2d} \int_{-d}^{+d} dz \mathbf{q}_{\text{VM}}^{\text{tan}}$$

is the value of $\mathbf{q}_{\text{VM}}^{\text{tan}}$ averaged over the slit.

If the walls of the channel are adiabatic, the viscomagnetic heat flux gives rise to a transverse temperature gradient. From the adiabatic condition $\mathbf{J}_q = 0$ and Eq. (4) one finds

$$(\nabla T)^{\text{tan}} = \frac{1}{\lambda} \left[\overline{\mathbf{q}_{\text{VM}}^{\text{tan}}} + \frac{1}{2d} (\mathbf{Q}(d) + \mathbf{Q}(-d)) \right]. \quad (5)$$

In the following sections this temperature gradient shall be calculated in lowest order of the inverse of the average pressure p_0 , viz. in the order p_0^{-1} . The surface heat flux in this approximation does not contain $(\nabla T)^{\text{tan}}$ itself.

3. Heat Flux in the Bulk Gas

In a non-equilibrium situation in a polyatomic gas, the rotational angular momenta $\mathbf{J}$ of the molecules, due to the nonspherical interaction, are partially aligned through collisions. For most gases, the (second rank) tensor polarization $\mathbf{a}^{(0200)}$,

$$\begin{aligned} a_{\mu\nu}^{(0200)} &= \langle \Phi_{\mu\nu}^{(0200)} \rangle, \\ \Phi_{\mu\nu}^{(0200)} &= \sqrt{\frac{15}{2}} \left(\langle J^2 (J^2 - \frac{3}{4}) \rangle_0 \right)^{-\frac{1}{2}} J_\mu J_\nu, \end{aligned} \quad (6)$$

and the (third rank) Kagan polarization $\underline{\mathbf{a}}^{(1200)}$,

$$a_{\lambda\mu\nu}^{(1200)} = \sqrt{\frac{m}{k_B T_0}} \langle c_\lambda \Phi_{\mu\nu}^{(0200)} \rangle, \quad (7)$$

are the dominant types of alignment in streaming and heat conducting gases⁷. The symbol $\langle \dots \rangle$ denotes a non-equilibrium average, $\langle \dots \rangle_0$ an equilibrium average, respectively; $\overline{[\dots]}$ refers to the symmetric traceless part of a tensor; m , $\mathbf{c}$ are the molecular mass and velocity, k_B is Boltzmann's constant and T_0 the equilibrium temperature. Application of the moment method²⁰ to the relevant kinetic equation due to Waldmann¹⁸ and Snider¹⁹ leads to a set of coupled transport relaxation equations for these polarizations and other physical quantities of interest, like the heat flux $\mathbf{q}$ and the friction pressure tensor $\overline{\mathbf{p}}$. So, the following ansatz for the bulk viscomagnetic heat flux is obtained:

$$\mathbf{q}_{VM} = \lambda \alpha T_0 p_0^{-1} (\delta - \Psi_1 \mathbf{\Lambda}(\varphi_{12})) \cdot \nabla p \quad (8)$$

$$+ \lambda \Psi_2 T_0 p_0^{-1} \nabla \cdot \sum_{n=-2}^{+2} (1 + i n \varphi_{12})^{-1} \mathbf{A}_n^{(02)}.$$

The first term in Eq. (8) is easily recognized as the generalization of Maxwell's viscous heat flux. The factor $\frac{2}{3}$ in Maxwell's work¹ is replaced by

$$\alpha = \frac{2}{5} \frac{\lambda_{tr}}{\lambda} (1 + \varepsilon), \quad (9)$$

i. e. in a polyatomic gas the translational part λ_{tr} of the total heat conductivity λ occurs. If values for the viscosity η and the bulk viscosity η_V are known, the ratio λ_{tr}/λ can be calculated from

$$\frac{\lambda_{tr}}{\lambda} = \left(\frac{15}{4} \frac{k_B}{m} \frac{\eta}{\lambda} + \frac{5}{6} c_r (1 - c_r) \frac{\eta}{\eta_V} \right) \quad (10)$$

$$\left/ \left(1 + \frac{5}{6} c_r \frac{\eta}{\eta_V} \right) \right.,$$

where $c_r = c_{rot}/(c_{rot} + \frac{3}{2} k_B)$ is the relative rotational specific heat. The correction term ε arises from the coupling of (eliminated) higher moments, e. g. $\mathbf{a}^{(2010)}$ and $\mathbf{a}^{(2001)}$. If the contributions due to inelastic collisions (i. e. change of rotational energy) are neglected, ε can be expressed²¹ in terms of Chapman-Cowling reduced Ω -integrals and is approximately given by

$$\varepsilon \approx \frac{\frac{7}{2} \Omega^{(2,2)*} - 4 \Omega^{(2,3)*}}{\frac{4}{12} \Omega^{(2,2)*} - 4 \Omega^{(2,3)*} + \frac{20}{9} \Omega^{(2,4)*}}. \quad (11)$$

Special values are $\varepsilon = 0$ for Maxwell molecules and $\varepsilon = -0.2049$ for hard spheres. The hard sphere

value is close to the numbers obtained (with different types of approximation) by Jones³ and by Burnett⁴. For a Lennard-Jones potential, ε is a function of the reduced temperature²²; typical values for gases like N_2 at room temperature are $\varepsilon \approx -0.1$.

A magnetic field $\mathbf{H} = H \mathbf{h}$ ($\mathbf{h} \cdot \mathbf{h} = 1$) influences the transport phenomena through the molecular magnetic moment $g_r \mu_N \mathbf{J}$ connected with the rotation, g_r is the rotational g -factor, μ_N the nuclear magneton. The second rank tensor $\mathbf{\Lambda}$ in Eq. (8),

$$\mathbf{\Lambda}_{\mu\nu}(\varphi_{12}) = h_\mu h_\nu f(\varphi_{12})$$

$$+ (\delta_{\mu\nu} - h_\mu h_\nu) \frac{1}{2} (f(\varphi_{12}) + 2 f(2\varphi_{12}))$$

$$+ \varepsilon_{\mu\lambda\nu} h_\lambda \frac{1}{2} (g(\varphi_{12}) + 2 g(2\varphi_{12})), \quad (12)$$

with

$$f(\varphi_{12}) = \varphi_{12}^2 / (1 + \varphi_{12}^2), \quad g(\varphi_{12}) = \varphi_{12} / (1 + \varphi_{12}^2),$$

contains the field direction $\mathbf{h}$ and depends on the field strength H via the precession angle $\varphi_{12} \propto H/p_0$. The precession angle

$$\varphi_{pq} \equiv \frac{\omega_H}{n_0 v_{rel} \mathfrak{S}(\frac{pq00}{pq00})}, \quad v_{rel} = (16 k_B T_0 / m)^{\frac{1}{2}}, \quad (13)$$

is the ratio of the precession frequency $\omega_H = (1/\hbar) g_r \mu_N H$ and a gas collision frequency; n_0 is the equilibrium density, $p_0 = n_0 k_B T_0$. Here, φ_{12} and φ_{02} will occur. With the help of $\mathbf{\Lambda}$, the heat conductivity tensor λ can be written as²³

$$\lambda = \lambda (\delta - \Psi^{(\lambda)} \mathbf{\Lambda}(\varphi_{12})), \quad (14)$$

where $\Psi^{(\lambda)}$ characterizes the relative size of the field effect*.

The second term in Eq. (8) is due to gradients of the tensor polarization. In the following, for convenience, the variables $\mathbf{A}^{(02)}$ and $\underline{\mathbf{A}}^{(12)}$ will be used instead of $\mathbf{a}^{(0200)}$ and $\underline{\mathbf{a}}^{(1200)}$:

$$\mathbf{A}^{(02)} = \sqrt{2} \mathfrak{S}(\frac{0200}{0200}) \mathfrak{S}^{-1}(\frac{2000}{0200}) p_0 \mathbf{a}^{(0200)}, \quad (15)$$

$$\underline{\mathbf{A}}^{(12)} = [2 \mathfrak{S}(\frac{0200}{0200}) \mathfrak{S}(\frac{1200}{1200})]^{\frac{1}{2}} \mathfrak{S}^{-1}(\frac{2000}{0200}) p_0 \underline{\mathbf{a}}^{(1200)}.$$

The spherical components $\mathbf{A}_n^{(02)} \equiv \mathcal{P}^{(n)} : \mathbf{A}^{(02)}$ of the "tensor polarization" $\mathbf{A}^{(02)} = \sum_{n=-2}^{+2} \mathbf{A}_n^{(02)}$ itself are determined by velocity gradients (see Section 4). Notice, that the fourth rank projection tensors²⁵

* For $\Psi^{(\lambda)}$ see e. g. Refs.^{23, 24}; with the constants used here, one has

$$\Psi^{(\lambda)} = 2 \lambda \left(\frac{k_B}{m} \eta \right)^{-1} \mathfrak{S}(\frac{1200}{1200}) \mathfrak{S}(\frac{0200}{0200}) \mathfrak{S}^{-2}(\frac{2000}{2000}) \Psi_2^2 / \Psi(\eta),$$

see Eqs. (16), (18) and (41) for Ψ_2 and $\Psi(\eta)$. The notation for the effective cross sections $\mathfrak{S}(\frac{pqrs}{pqrs})$ is the same as in Reference²⁴.

$\mathcal{P}^{(n)}$ depend on the field direction $\mathbf{h}$. The relative size of the two field dependent terms in the bulk viscomagnetic heat flux is determined by the values of the two constants Ψ_1 and Ψ_2 which can be expressed in terms of effective cross sections in the following way:

$$\Psi_1 = f_1 \Psi_0, \Psi_2 = \frac{2}{3} (\lambda_{tr}/\lambda) \Psi_0, \quad (16)$$

with

$$f_1 = \sqrt{\frac{4}{5}} \frac{\mathfrak{S}_{(0200)}^{(2000)}}{\mathfrak{S}_{(0200)}^{(2000)}} \cdot [-\mathfrak{S}_{(1200)}^{(1010)} \mathfrak{S}_{(1001)}^{(1001)} + \mathfrak{S}_{(1200)}^{(1001)} \mathfrak{S}_{(1001)}^{(1010)}] \cdot [\mathfrak{S}_{(1010)}^{(1010)} \mathfrak{S}_{(1001)}^{(1001)} - \mathfrak{S}_{(1001)}^{(1010)}]^{-1}, \quad (17)$$

and

$$\Psi_0 = \sqrt{\frac{5}{4}} \frac{\mathfrak{S}_{(0200)}^{(2000)}}{\mathfrak{S}_{(0200)}^{(2000)} \mathfrak{S}_{(1200)}^{(1001)}} \left[-\mathfrak{S}_{(1200)}^{(1010)} + \mathfrak{S}_{(1001)}^{(1001)} \frac{\mathfrak{S}_{(1010)}^{(1010)} + \sqrt{\frac{2}{5}} \frac{c_{rot}}{k_B} \mathfrak{S}_{(1010)}^{(1010)}}{\mathfrak{S}_{(1001)}^{(1001)} + \sqrt{\frac{2}{5}} \frac{c_{rot}}{k_B} \mathfrak{S}_{(1010)}^{(1010)}} \right]. \quad (18)$$

According to Eq. (5), the transverse temperature gradient is determined by the averaged viscomagnetic heat flux. With the assumption that the tensor polarization depends on z only, the integration of Eq. (8) gives

$$\bar{\mathbf{q}}_{VM} = \lambda \alpha T_0 p_0^{-1} [\bar{\nabla} p - \Psi_1 \mathbf{\Lambda}(\varphi_{12}) \cdot \nabla p^{(0)}] + \lambda \Psi_2 T_0 p_0^{-1} \frac{1}{2d} \mathbf{e}_z \cdot \sum_{n=-2}^{+2} (1 + i n \varphi_{12})^{-1} \cdot [\mathbf{A}_n^{(02)}(d) - \mathbf{A}_n^{(02)}(-d)]. \quad (19)$$

For the evaluation of the first term in Eq. (19), the magnetic field depend pressure gradient in the streaming gas has to be known. In the second term however, the field free value $\nabla p^{(0)}$ can be inserted, if contributions of the order $\Psi_1 \Psi^{(n)}$ are neglected. The third term in Eq. (19) is given by the boundary values of the tensor polarization, hence boundary layer phenomena will be of importance for the average viscomagnetic heat flux in the bulk.

4. Boundary Conditions, Surface Heat Flux

For the calculation of the transverse temperature gradient, boundary conditions for the tensor polarization and an ansatz for the surface heat flux are needed. Both are derived by the thermodynamical method introduced by Waldmann¹⁴. Starting point is the entropy production at the interface between

two media, i. e. the gas and a solid. According to the principles of non-equilibrium thermodynamics, the surface entropy production is transformed into a sum of products of independent “fluxes” and “forces”, using the local conservation laws at the interface. In particular, the energy law,

$$(\mathbf{q} - \mathbf{q}_s) \cdot \mathbf{n} + \mathbf{k} \cdot \mathbf{v} = \nabla^\sigma \cdot \mathbf{Q}, \quad (20)$$

connects the discontinuity of the normal heat fluxes in the gas ($\mathbf{q}$) and in the solid ($\mathbf{q}_s$), and the work term ($\mathbf{k} \cdot \mathbf{v}$) with the divergence of the surface heat flux $\mathbf{Q}$. Here, $\mathbf{n}$ is the outward unit normal of the gas at the interface, $\mathbf{k} = \mathbf{p} \cdot \mathbf{n}$ is the force the gas exerts per unit area of the solid, and ∇^σ denotes a two dimensional surface derivative. Linear relations set up between “fluxes” and “forces” supply the boundary conditions^{14, 15}. So, for the surface heat flux an ansatz is obtained which is coupled with two other vectorial conditions¹⁷, viz. the velocity slip condition and the boundary condition for the tensor polarization components $\mathbf{a}^{(0200)} \equiv (\mathbf{a}^{(0200)} \cdot \mathbf{n})^{\tan}$:

$$-T_0^{-1} \mathbf{Q} = C_h l \lambda T_0^{-1} \nabla^{\tan} \frac{1}{2} (\tilde{T} + T_s) + C_{hm} \lambda p_0^{-1} \mathbf{k}^{\tan} + C_{ha} \lambda \mathbf{a}^{(1200)}, \quad (21)$$

$$\tilde{\mathbf{v}}^{\tan} = C_{mh} \lambda p_0^{-1} \nabla^{\tan} \frac{1}{2} (\tilde{T} + T_s) + C_m l \eta^{-1} \mathbf{k}^{\tan} + C_{ma} c_0 \mathbf{a}^{(1200)}, \quad (22)$$

$$\mathbf{a}^{(0200)} = C_{ah} \lambda (2 p_0 c_0)^{-1} \nabla^{\tan} \frac{1}{2} (\tilde{T} + T_s) + C_{am} (2 p_0)^{-1} + C_a \mathbf{a}^{(1200)}. \quad (23)$$

Here, $\mathbf{a}^{(1200)}$ stands for $a_v^{(1200)} = (a_{\lambda\mu\nu}^{(1200)} n_\lambda n_\mu)^{\tan}$, and l is a viscosity characteristic free path,

$$l = \eta c_0 p_0^{-1}, \quad c_0 = (k_B T_0/m)^{\frac{1}{2}}.$$

Typically for a consistent treatment of the rarefied gas, the tilded velocity $\tilde{\mathbf{v}}$ and tilded temperature $\tilde{T}$ occur in the boundary conditions instead of the actual velocity and temperature^{15, 16},

$$\tilde{\mathbf{v}} = \mathbf{v} + \frac{2}{3} (1 + \varepsilon) p_0^{-1} \mathbf{q}_{tr},$$

$$\tilde{T} = T + \frac{2}{3} (1 + \varepsilon) c_t T_0 p_0^{-1} \mathbf{n} \cdot \bar{\mathbf{p}} \cdot \mathbf{n},$$

where $\mathbf{q}_{tr}$ is the translational heat flux, and $c_t = \frac{3}{2} k_B / (c_{rot} + \frac{3}{2} k_B)$ is the relative translational specific heat. The phenomenological coefficients $C_{..}$ are subject to two restrictive conditions. One is a consequence of the second law of thermodynamics, requiring a positive surface entropy production, i. e. a positive definite $C_{..}$ -matrix:

$$C_k > 0, \quad (24)$$

$$C_k C_l - \xi^{(kl)} \left(\frac{C_{kl} + C_{lk}}{2} \right)^2 > 0, \quad l, k = a, h, m. \quad (25)$$

The factors $\xi^{(kl)} = \xi^{(lk)} > 0$,

$$\xi^{(\text{mh})} = \lambda \left(\frac{k_B}{m} \eta \right)^{-1}, \quad \xi^{(\text{ma})} = \frac{1}{2}, \quad \xi^{(\text{ha})} = \frac{1}{2} \xi^{(\text{mh})}, \quad (26)$$

enter automatically through the use of the dimensionless slip numbers $C_{\cdot}$ instead of the coefficients $L_{\cdot}^{(V)}$ of Ref. 17. A second requirement stems from Onsager relations:

$$C_{\text{ma}} = -C_{\text{am}}, \quad C_{\text{ha}} = -C_{\text{ah}}, \quad (27)$$

$$C_{\text{mh}} = C_{\text{hm}} \equiv \sigma. \quad (28)$$

With the anti-symmetry (27), Eq. (25) yields nothing new beyond Eq. (24), whereas the combination of the symmetry (28) with (25) gives a restriction¹⁶ for the absolute value of the thermal slip number σ ,

$$C_{\text{m}} C_{\text{h}} - \xi^{(\text{mh})} \sigma^2 > 0.$$

Next, the general boundary conditions (21)–(23) are discussed briefly. According to Eq. (21), the surface heat flux $\mathbf{Q}$ is determined by temperature gradients along the wall, by tangential stresses and by the Kagan polarization component $\mathbf{a}^{(1200)}$. For the tangential velocity in Eq. (22), there are several types of slip: two sources of thermal slip, due to a tangential heat flux (hidden in $\tilde{\mathbf{v}}$) and to tangential temperature gradients, the mechanical slip ($\propto C_{\text{m}} \mathbf{k}^{\text{tan}}$) and the thermomagnetic slip^{10, 17} ($\propto C_{\text{ma}} \mathbf{a}^{(1200)}$). The boundary value of the tensor polarization component $\mathbf{a}^{(0200)}$ results from two production mechanisms, characterized by C_{ah} and C_{am} , and an accommodation process ($\propto C_{\text{a}} \mathbf{a}^{(1200)}$). Values for the six independent slip numbers can be determined in the following way: C_{m} is found from deviations of the frictional force from Stokes' formula, σ from thermal transpiration in the high pressure limit, and then C_{h} is obtained from the analysis of thermal force data¹⁶. Numerical values for C_{a} and C_{am} are derived from Knudsen corrections for the field effect of viscosity²⁶ and from thermomagnetic pressure difference data¹². The only coefficient not yet known, $C_{\text{ha}} = -C_{\text{ah}}$, can hopefully be inferred from viscomagnetic heat flux experiments.

Now, the general boundary conditions (21) and (23) shall be specialized to the case of a flowing gas. For the approximate calculation of the polarizations, the first term in Eq. (23) can be neglected since it is nonzero only in the presence of a field, thus it is an order of magnitude smaller than the two remaining terms. Similarly, the tangential force $\mathbf{k}^{\text{tan}}$ can be replaced by its "isotropic" (field free)

value $\mathbf{k}_0^{\text{tan}}$. In addition to Eq. (23), there are also boundary conditions for the surface scalar part $\mathbf{n} \mathbf{n} : \mathbf{a}^{(0200)}$ and the surface tensor part $\mathbf{a}^{(0200)}$ of the tensor polarization¹⁷. However it is assumed that they are related by the same C_{a} to the corresponding components of the normal Kagan polarization $n_{\lambda} a_{\lambda\mu\nu}^{(1200)}$. So, for the present situation one has

$$\mathbf{n} \mathbf{n} : \mathbf{a}^{(0200)} = C_{\text{a}} \mathbf{n} \mathbf{n} : (\mathbf{n} \cdot \mathbf{a}^{(1200)}), \\ \mathbf{a}^{(0200)} = C_{\text{a}} \mathbf{a}^{(1200)}.$$

Thus, a boundary condition for the total tensor $\mathbf{a}^{(0200)}$ can be stated; for the variables $\mathbf{A}^{(02)}$ and $\mathbf{A}^{(12)}$ introduced in Eq. (15), it reads

$$\mathbf{A}^{(02)} = \tilde{C}_{\text{am}} 2 \overline{\mathbf{k}_0^{\text{tan}}} \mathbf{n} + \tilde{C}_{\text{a}} \mathbf{n} \cdot \mathbf{A}^{(12)}, \quad (29)$$

where $\tilde{C}_{\text{am}}$, $\tilde{C}_{\text{a}}$ are related to C_{am} , C_{a} by

$$C_{\text{am}} = \sqrt{2} \frac{\mathcal{S}_{(0200)}^{(2000)}}{\mathcal{S}_{(0200)}^{(0200)}} \tilde{C}_{\text{am}}, \\ C_{\text{a}} = [\mathcal{S}_{(1200)}^{(1200)} / \mathcal{S}_{(0200)}^{(0200)}]^{1/2} \tilde{C}_{\text{a}}. \quad (30)$$

In the ansatz (21) for the surface flux, the first term can be neglected since it is of order p_0^{-2} and here only terms of the order p_0^{-1} shall be retained. Using $\tilde{C}_{\text{ha}}$ instead of C_{ha} ,

$$C_{\text{ha}} = \sqrt{2} [\mathcal{S}_{(0200)}^{(0200)} / \mathcal{S}_{(1200)}^{(1200)}]^{1/2} \mathcal{S}_{(0200)}^{-1(2000)} \Psi_2 \tilde{C}_{\text{ha}}, \quad (31)$$

one finds

$$\mathbf{Q} = -\lambda T_0 p_0^{-1} (\sigma \mathbf{k} + \tilde{C}_{\text{ha}} \Psi_2 \mathbf{n} \mathbf{n} : \mathbf{A}^{(12)})^{\text{tan}}. \quad (32)$$

For the explicit calculation of $\mathbf{Q}$, the magnetic field dependent force $\mathbf{k}^{\text{tan}}$ (i. e. the pressure gradient) and the Kagan polarization have to be determined.

5. Tensor Polarization in a Streaming Gas

By insertion of the ansatz for the Kagan polarization into the moment equation for the tensor polarization, second order differential equations for the spherical components $\mathbf{A}_n^{(02)}$ are obtained¹⁷. For a streaming gas with $\mathbf{A}_n^{(02)} = \mathbf{A}_n^{(02)}(z)$ they are

$$\left(1 - L_n^2 \frac{d^2}{dz^2} \right) \mathbf{A}_n^{(02)} = \frac{2 \eta}{1 + i n \varphi_{02}} (\overline{\nabla \mathbf{v}})_n \\ = \frac{2 z}{1 + i n \varphi_{02}} (\overline{\mathbf{e}_z \nabla p^{(0)}})_n, \\ n = 0, \pm 1, \pm 2. \quad (33)$$

Here, the velocity gradient has been replaced by its field-free value for plane Poiseuille flow. The mag-

netic field dependent quantity L_n has the dimension of a length

$$L_n = L_0 [(1 + i n \varphi_{02}) (1 + i n \varphi_{12})]^{-\frac{1}{2}}, \quad (34a)$$

where the free path L_0 ,

$$L_0 = n_0^{-1} v_{\text{rel}}^{-1} (k_B T_0 / m)^{\frac{1}{2}} [\zeta(\frac{0200}{0200}) \zeta(\frac{1200}{1200})]^{-\frac{1}{2}} \\ = n_0^{-1} \frac{1}{4} \sqrt{\pi} [\zeta(\frac{0200}{0200}) \zeta(\frac{1200}{1200})]^{-\frac{1}{2}}, \quad (34b)$$

somehow characterizes the thickness of the layer in which the transition from the boundary value to the bulk value of tensor polarization takes place.

The solution of Eq. (33),

$$\mathbf{A}_n^{(02)} = \frac{2z}{1 + i n \varphi_{02}} \overline{(\mathbf{e}_z \nabla p^{(0)})}_n + \mathbf{B}_n, \quad (35)$$

is the sum of the bulk value of tensor polarization, first term in Eq. (35), and of a "boundary layer" term $\mathbf{B}_n$ which satisfies the homogeneous differential equation

$$\left(1 - L_n^2 \frac{d^2}{dz^2}\right) \mathbf{B}_n = \mathbf{0}.$$

Consequently, $\mathbf{B}_n$ will be the superposition of a $\sinh(z/L_n)$ and a $\cosh(z/L_n)$ term. For the inhomogeneity in the boundary condition (29) one has, for $z = \pm d$, $\mathbf{k}_0^{\text{tan}} = -d \nabla p^{(0)}$ and $\mathbf{n} = \pm \mathbf{e}_z$, so that $\mathbf{A}_n^{(02)}$ has to be an odd function of z . Hence, the solution for $\mathbf{B}_n$ can be written in the form

$$\mathbf{B}_n = \beta_n \frac{\sinh(z/L_n)}{\sinh(d/L_n)} 2d \overline{(\mathbf{e}_z \nabla p^{(0)})}_n, \quad (36)$$

and with the ansatz for the Kagan polarization,

$$\mathbf{e}_z \cdot \mathbf{A}^{(12)} = -L_0 \sum_{n=-\infty}^{+\infty} \frac{1}{1 + i n \varphi_{12}} \cdot \left[\frac{d}{dz} \mathbf{A}^{(02)} + (1 + \varepsilon) f_1 \overline{\mathbf{e}_z \nabla p^{(0)}} \right]_n, \quad (37)$$

the constant β_n is found from the boundary condition (29) as

$$\beta_n = - \left(\frac{1}{1 + i n \varphi_{02}} - \tilde{C}_{\text{ma}} \right) / \left(1 + \tilde{C}_a \sqrt{\frac{1 + i n \varphi_{02}}{1 + i n \varphi_{12}}} \right). \quad (38)$$

This actually is the simplest approximation for β_n , neglecting terms of the order L_0/d and replacing $\cosh(d/L_n)$ by 1.

According to Eq. (36), the surface contribution $\mathbf{B}_n$ to the total tensor polarization profile is different from zero only in a boundary layer of thickness L_0 . Further away from the wall, only the bulk contribution, first term in Eq. (35), remains. For the value

of the tensor polarization at the plates,

$$\mathbf{A}_n^{(02)}(d) = -\mathbf{A}_n^{(02)}(-d) \quad (39) \\ = \left[\frac{1}{1 + i n \varphi_{02}} + \beta_n \right] 2d \overline{(\mathbf{e}_z \nabla p^{(0)})}_n$$

both contributions may be of the same order of magnitude. In Levi, McCourt and Beenakker's approximation⁶, only the bulk term [first term in Eq. (39)] was taken into account. Neglecting the surface contribution is, from the present point of view, equivalent to assuming large $\tilde{C}_a$ -values, $\tilde{C}_a \gg 1$ and $\tilde{C}_a \gg |\tilde{C}_{\text{ma}}|$; realistic values of $\tilde{C}_a$, however, are of the order of unity^{12, 26}. In order to get some feeling for the relation between the size of the wall effects and the magnitude of the surface parameter $\tilde{C}_{\text{ma}}$, two special cases of Eq. (39) are discussed now. For zero field, the "surface contribution" $\beta_0 = (\tilde{C}_{\text{ma}} - 1)/(\tilde{C}_a + 1)$ has to be compared with the "bulk value" of 1. With $\tilde{C}_{\text{ma}} < 1$, one gets $\beta_0 < 0$, i. e. different types of tensor polarization are produced in the boundary layer and in the bulk gas. For high fields, the bulk contribution vanishes, whereas the surface contribution is proportional to

$$\beta_n = \tilde{C}_{\text{ma}} / (1 + \tilde{C}_a \sqrt{\zeta(\frac{1200}{1200}) / \zeta(\frac{0200}{0200})}).$$

Hence, the sign of $\tilde{C}_{\text{ma}}$ determines the type of tensor polarization which survives near the wall. More details about the tensor polarization profile (35), (36) can be found in Reference²⁶.

6. The Magnetic Field Dependent Pressure Gradient

The complete expression for the pressure gradient is obtained from the solution for plane Poiseuille flow in a magnetic field²⁶. In the limit $L_0/d = 0$ however, the expression for $\partial p / \partial y$ can easily be derived, using the following arguments. Basically, one assumes that ∇p is a function of z only, then $\partial p / \partial x$ and $\partial p / \partial y$ must be constant. By insertion of the ansatz for the friction pressure tensor

$$\bar{\mathbf{p}} = -2\eta^{[0]} (\nabla \mathbf{v} + \frac{2}{5}(1 + \varepsilon) p_0^{-1} \nabla q_{\text{tr}}) - \Psi^{(\eta)} \mathbf{A}^{(02)}$$

into the steady state equation for momentum conservation,

$$\nabla p = -\nabla \cdot \bar{\mathbf{p}},$$

one readily finds, using $\nabla \cdot \mathbf{v} = 0$, $\nabla \cdot \mathbf{q}_{\text{tr}} = 0$,

$$\nabla p = \eta^{[0]} \nabla \cdot \nabla (\mathbf{v} + \frac{2}{5}(1 + \varepsilon) p_0^{-1} \mathbf{q}_{\text{tr}}) \\ + \Psi^{(\eta)} \nabla \cdot \mathbf{A}^{(02)}. \quad (40)$$

Here, $\eta^{[0]}$ denotes the "isotropic" value of shear viscosity, and the quantity $\Psi^{(\eta)}$,

$$\Psi^{(\eta)} = \frac{\mathfrak{S}^2_{(0200)}(2000)}{\mathfrak{S}_{(2000)}(2000) \mathfrak{S}_{(0200)}(0200)}, \quad (41)$$

characterizes the relative magnitude of the viscosity field effect. The last term on the right hand side of Eq. (40) gives a constant and an exponentially varying contribution which vanishes in the bulk gas. Since the second term reduces to third order derivatives of $\mathbf{A}^{(02)}$, only the exponentials (from $\mathbf{B}_n$) remain. Finally let us look at the y -component of the flow velocity; v_y will be the sum of a parabolic profile and of terms decaying exponentially with increasing distance from the walls. By the flow in y -direction, no gas shall leave the channel, so that $\bar{v}_y = 0$ must be required. Consequently, in the bulk gas, v_y is of order L_0/d . Therefore, $\nabla \cdot \nabla v_y$ contains constant terms of the order L_0/d and exponential terms. As a conclusion one finds that the constant transverse pressure gradient $\partial p / \partial y$, for $L_0/d = 0$, is determined by the constant terms in $\nabla \cdot \mathbf{A}^{(02)}$ only, hence from Eqs. (35), (40)

$$\frac{\partial p}{\partial y} = \Psi^{(\eta)} \sum_{n=-2}^{+2} \frac{2}{1 + i n \varphi_{02}} (\mathbf{e}_y \mathbf{e}_z) : [\overline{\mathbf{e}_z \nabla p^{(0)}}]_n. \quad (42)$$

7. The Transverse Temperature Gradient

By insertion of the results (39) and (42) for $\mathbf{A}^{(02)}(d)$ and $\partial p / \partial y$ into Eq. (19), the y -component of the average viscomagnetic heat flux in the bulk is obtained as

$$\begin{aligned} \mathbf{e}_y \cdot \bar{\mathbf{q}}_{\text{VM}} = & \lambda T_0 p_0^{-1} \left\{ -\alpha \Psi_1 \mathbf{e}_y \cdot \mathbf{A}(\varphi_{12}) \cdot \nabla p^{(0)} \right. \\ & + \sum_{n=-2}^{+2} 2(\mathbf{e}_y \mathbf{e}_z) : [\overline{\mathbf{e}_z \nabla p^{(0)}}]_n \\ & \cdot \left[\frac{\alpha \Psi^{(\eta)}}{1 + i n \varphi_{02}} + \frac{\Psi_2}{1 + i n \varphi_{12}} \left(\frac{1}{1 + i n \varphi_{02}} + \beta_n \right) \right] \Bigg\}. \end{aligned} \quad (43)$$

The term with the β_n is the surface contribution, whereas the three remaining terms constitute the bulk contribution already given by Levi, McCourt and Beenakker⁶.

In addition to the bulk heat flux $\mathbf{q}_{\text{VM}}$ there is also the surface heat flux $\mathbf{Q}$ which is given by the boundary condition (32). For the flow between infinite parallel plates, the first term in Eq. (32) is simply found from Eq. (42) and from

$$\mathbf{k}^{\text{tan}}(d) = \mathbf{k}^{\text{tan}}(-d) = -d (\nabla p)^{\text{tan}}.$$

The second term in Eq. (32) is calculated with the help of the ansatz (37) for the Kagan polarization. Here, in lowest order in L_0/d , only the exponential $\mathbf{B}_n$ -term in $\mathbf{A}^{(02)}$ has to be retained in Eq. (37); in the same approximation, $\coth(d/L_n)$ is replaced by 1. The result for the y -component of the surface heat flux is

$$\begin{aligned} Q_y(d) = & Q_y(-d) \\ = & d \lambda T_0 p_0^{-1} \sum_{n=-2}^{+2} 2(\mathbf{e}_y \mathbf{e}_z) : (\overline{\mathbf{e}_z \nabla p^{(0)}})_n \\ & \cdot \left[\frac{\sigma \Psi^{(\eta)}}{1 + i n \varphi_{02}} + \tilde{C}_{\text{ha}} \Psi_2 \beta_n \right] \sqrt{\frac{1 + i n \varphi_{02}}{1 + i n \varphi_{12}}}. \end{aligned} \quad (44)$$

Now, with Eqs. (43), (44) the general expression for the transverse temperature gradient $\partial T / \partial y$ can be derived from Equation (5). Of practical interest⁹ in particular is the case where the magnetic field varies in the plane perpendicular to the field-free flow direction $\mathbf{e}_x$,

$$\mathbf{h} = \cos \chi \mathbf{e}_y + \sin \chi \mathbf{e}_z. \quad (45)$$

Using Eq. (38) for β_n , and evaluating the dependence on the field direction [with Eq. (12) and the Appendix of Ref. 26] with $\nabla p^{(0)} = \mathbf{e}_x (\partial p / \partial x)_0$ leads to the explicit result:

$$\frac{\partial T}{\partial y} = -\frac{T_0}{p_0} \left(\frac{\partial p}{\partial x} \right)_0 \sin \chi [\delta + \Delta_1 + \cos^2 \chi (\Delta_2 - 2 \Delta_1)], \quad (46)$$

where

$$\delta = \alpha \Psi_1 \frac{1}{2} (g(\varphi_{12}) + 2 g(2 \varphi_{12})), \quad (47)$$

$$\begin{aligned} \Delta_n = & \left((\sigma + \alpha) \Psi^{(\eta)} + \frac{r}{r-1} \Psi_2 \right) g(n \varphi_{02}) \\ & - \frac{1}{r-1} \Psi_2 g(n \varphi_{12}) \\ & + \Psi_2 \operatorname{Im} \left[\frac{1}{1 + i n \varphi_{02}} - \tilde{C}_{\text{ma}} \right. \\ & \cdot \left. \left(\frac{1}{1 + i n \varphi_{12}} + \tilde{C}_{\text{ha}} \sqrt{\frac{1 + i n \varphi_{02}}{1 + i n \varphi_{12}}} \right) \right], \quad n = 1, 2, \end{aligned} \quad (48)$$

and r is an abbreviation for the ratio

$$r = \mathfrak{S}_{(1200)}(1200) / \mathfrak{S}_{(0200)}(0200).$$

According to Eq. (46), the transverse temperature gradient is inversely proportional to the average gas pressure p_0 , and is proportional to the field-

free pressure gradient $(\partial p / \partial x)_0$. If the magnetic field is perpendicular to the plates (and to the field-free flow direction $\mathbf{e}_x$), $\chi = 90^\circ$, the temperature gradient as a function of H/p_0 reaches its largest absolute value if $|\varphi_{02}|$, $|\varphi_{12}|$ are about unity. With decreasing angle χ , this maximum decreases about like $\sin \chi$ and shifts to lower H/p_0 -values, since for $\chi < 90^\circ$ the double frequencies $2\varphi_{02}$, $2\varphi_{12}$ (through Δ_2) come into play. The transverse pressure gradient [according to Eq. (42)],

$$\frac{\partial p}{\partial y} = - \left(\frac{\partial p}{\partial x} \right)_0 \Psi^{(\eta)} \sin \chi \cdot [g(\varphi_{02}) + \cos^2 \chi (g(2\varphi_{02}) - 2g(\varphi_{02}))], \quad (49)$$

shows the same type of angular dependence. Notice, that the first terms in Eq. (49) and in Δ_1 have the magnetic field dependence of the viscosity coefficient η_5 ,

$$\eta_5 = -\eta \Psi^{(\eta)} g(\varphi_{02}),$$

and that the quantity δ is proportional to the transverse heat conductivity coefficient

$$\lambda^{\text{transv}} = -\lambda \Psi^{(\lambda)} \frac{1}{2} (g(\varphi_{12}) + 2g(2\varphi_{12})).$$

For $\chi = 90^\circ$, the maximum of $|\partial T / \partial y|$ as a function of H/p_0 will therefore occur somewhere between that of $|\eta_5|$ and of $|\lambda^{\text{transv}}|$.

Now let us look more closely into some details of Eqs. (47) and (48). The δ -term and the second, third and fourth terms in Δ_n are the bulk contributions already given by Levi, McCourt and Beenakker⁶. Their result is obtained from our general expression (47), (48) with $\varepsilon = 0$ [i. e. $\alpha = \frac{2}{3} \lambda_{\text{tr}} / \lambda$, see Eq. (9)], $\sigma = 0$ and $\tilde{C}_{\text{ma}} = 0 = \tilde{C}_{\text{ha}}$, $\tilde{C}_a \gg 1$. The remaining terms in Δ_n are the surface contributions, two of them originating from the surface heat flux (proportional to σ and $\tilde{C}_{\text{ha}}$), and the others (involving $\tilde{C}_a$ and $\tilde{C}_{\text{ma}}$) from the wall effects for the tensor polarization. The relative importance of the surface contributions is determined by the values for

the bulk constants Ψ_1 , Ψ_2 , $\Psi^{(\eta)}$, and by those for the surface parameters $\tilde{C}_a$, $\tilde{C}_{\text{ma}}$, $\tilde{C}_{\text{ha}}$. The constants Ψ_1 , Ψ_2 and $\Psi^{(\eta)}$ can be calculated from Eqs. (16) – (18) and (41) with the values for effective cross sections derived from sound absorption measurements (rotational relaxation, η_V) and from data on the field effects of viscosity and thermal conductivity, see e. g. the values compiled in Reference²⁷. Thus, the bulk contributions to $\partial T / \partial y$ are known, see Table I. The two surface parameters $\tilde{C}_a$ and $\tilde{C}_{\text{ma}}$ determine the magnitude of the Knudsen corrections for the viscosity field effect²⁶ and for flow birefringence²⁸. Furthermore, they enter into the theoretical expression for the thermomagnetic pressure difference¹³. From this information, sets of $\tilde{C}_a$ and $\tilde{C}_{\text{ma}}$ can be found for these gases¹².

In the next section, it will be shown how the value of the yet unknown parameter $\tilde{C}_{\text{ha}}$ influences the magnitude of the transverse temperature gradient $\partial T / \partial y$, the shape of the curve as a function of H/p_0 , and the dependence on the angle χ .

8. Quantitative Results

For the following discussions and graphs it is convenient to use the reduced temperature gradient τ instead of $\partial T / \partial y$ itself:

$$\tau \equiv - \frac{p_0}{T_0} \frac{\partial T}{\partial y} / \left(\frac{\partial p}{\partial x} \right)_0 = \sin \chi [\delta + \Delta_1 + \cos^2 \chi (\Delta_2 - 2\Delta_1)]. \quad (50)$$

The dimensionless quantity τ is independent of the average pressure p_0 and of the field-free pressure gradient $(\partial p / \partial x)_0$.

First let us consider the case of $\chi = 90^\circ$ where the magnetic field is perpendicular to the plates and to the field-free flow direction $\mathbf{e}_x$. According to Eqs. (47), (48), the bulk contribution τ^b to the total

Table I. Values for the gas parameters ε , $\lambda_{\text{tr}} / \lambda$, α , Ψ_1 , Ψ_2 , $\Psi^{(\eta)}$ calculated from Eqs. (9) – (11), (16) – (18) with the help of Ref.²² and with the experimental data listed in Ref.²⁷. The bulk contribution τ^b to the reduced temperature gradient (for $\chi = 90^\circ$) has its maximum value $|\tau^b|_{\text{max}}$ at $(H/p_0)_{\text{max}}^b$, see Eqs. (50), (51).

	ε	$\frac{\lambda_{\text{tr}}}{\lambda}$	α	$10^3 * \Psi_1$	$10^3 * \Psi_2$	$10^3 * \Psi^{(\eta)}$	$(H/p_0)_{\text{max}}^b$ kOe/torr	$10^3 * \tau^b _{\text{max}}$
HD	−0.14	0.715	0.246	0.50	1.61	1.97	0.47	1.20
CH ₄	−0.07	0.609	0.226	0.87	0.34	0.80	5.26	0.44
CO	−0.09	0.720	0.262	4.49	1.50	3.52	4.39	2.23
N ₂	−0.11	0.715	0.254	3.69	1.53	2.72	3.49	1.94

reduced temperature gradient τ is given by

$$\tau^b = \left(\alpha \Psi^{(\eta)} + \frac{r}{r-1} \Psi_2 \right) g(\varphi_{02}) \quad (51)$$

$$- \frac{1}{r-1} \Psi_2 g(\varphi_{12}) + \delta.$$

The total effect also contains the surface contribution τ^s , $\tau = \tau^b + \tau^s$. For the gases HD, CH₄, CO and N₂, the bulk contribution has the same sign as the rotational g -factor, viz. $\tau^b > 0$ for HD, CH₄, and $\tau^b < 0$ for CO, N₂. The maximum value of $|\tau^b|$ is of the order 10^{-3} , see Table I. For N₂ and HD, the bulk contribution is plotted as a function of H/p_0 in Figs. 1, 2, together with the surface contributions and the total effect for a set of $\tilde{C}_a$, $\tilde{C}_{ma}$ (see Ref. ¹²) and three values of $\tilde{C}_{ha}$. In the case of $\tilde{C}_{ha} = 0$, the surface contributions subtract from the bulk contributions, viz. for N₂ over the whole H/p_0 -range, and for HD at lower H/p_0 -values (i.e. $H/p_0 \leq 2$ kOe/torr). With large negative values of $\tilde{C}_{ha}$, however, the surface contributions add to the bulk contributions, e.g. for $\tilde{C}_{ha} = -3.0$ in the case of N₂ (see Figure 1). This behaviour can be qualitatively understood from the discussion of the two limits $\varphi_{pq} \rightarrow 0$ and $\varphi_{pq} \rightarrow \infty$ of Eqs. (47) and (48).

For $|\varphi_{pq}| \ll 1$ one finds

$$\tau^b \approx \varphi_{12} \left\{ \frac{5}{2} \alpha \Psi_1 + (r+1) \Psi_2 + \alpha r \Psi^{(\eta)} \right\}, \quad (52a)$$

$$\tau^s \approx \varphi_{12} \left\{ \sigma r \Psi^{(\eta)} - \frac{\Psi_2}{(\tilde{C}_a + 1)^2} \cdot \left[r+1 + \frac{3r+1}{2} \tilde{C}_a - \tilde{C}_{ma} \left(1 + \frac{r+1}{2} \tilde{C}_a \right) + \tilde{C}_{ha} \left(\frac{r+1}{2} + r \tilde{C}_a + \frac{r-1}{2} \tilde{C}_{ma} \right) \right] \right\}. \quad (52b)$$

In the case of $\tilde{C}_{ha} = 0$, τ^s would have the same sign as τ^b only for large positive values of $\tilde{C}_{ma}$ ($\tilde{C}_a > 0$, and usually $\sigma < 0$); e.g. for N₂ ($r = 2.0$) if $\tilde{C}_{ma} > 2.6$, and for HD ($r = 6.1$) if $\tilde{C}_{ma} > 0$. With the $\tilde{C}_{ma}$ -values used in Figs. 1, 2, $\tau^s \varphi_{12}^{-1}$ is negative, whereas $\tau^b \varphi_{12}^{-1}$ is positive. Negative values of $\tilde{C}_{ha}$ increase $\tau^s \varphi_{12}^{-1}$ and make it finally positive, viz. for $\tilde{C}_{ha} < -2.43$ in the case of N₂, and for $\tilde{C}_{ha} < -1.51$ in the case of HD.

In the limit $|\varphi_{pq}| \gg 1$ on the other hand, one finds

$$\tau^b \approx \varphi_{12}^{-1} \alpha (\Psi_1 + r^{-1} \Psi^{(\eta)}), \quad (53a)$$

$$\tau^s \approx \varphi_{12}^{-1} [r^{-1} \sigma \Psi^{(\eta)} + \Psi_2 (\tilde{C}_{ma} - \sqrt{r^{-1}} \tilde{C}_{ha}) / (1 + \sqrt{r} \tilde{C}_a)]. \quad (53b)$$

Here $\tau^b \varphi_{12}$ is positive, whereas for $\tilde{C}_{ha} = 0$ and $\tilde{C}_{ma} < 0$ (e.g. N₂) $\tau^s \varphi_{12}$ is negative; for HD with $\tilde{C}_{ma} = 0.4$ and $\tilde{C}_{ha} = 0$, $\tau^s \varphi_{12}$ is positive (see Figs. 1 and 2). Again, negative $\tilde{C}_{ha}$ -values increase $\tau^s \varphi_{12}$, e.g. for N₂, $\tau^s \varphi_{12}$ gets positive if $\tilde{C}_{ha} < -0.62$.

With these considerations one can qualitatively understand Figs. 1 and 2. Furthermore, one can see that the total effect (τ) increases with increasing $\tilde{C}_{ma}$ and decreasing $\tilde{C}_{ha}$. The absolute value of the surface contribution τ^s decreases with increasing $\tilde{C}_a$. The dependence of τ on $\tilde{C}_a$ and $\tilde{C}_{ma}$ is in accordance with that of the tensor polarization discussed in Section 5.

From the preceding discussion it is obvious that for $\tilde{C}_{ha} \neq 0$ the shape of the total curve $\tau = \tau(H/p_0)$ will change considerably compared with the case of $\tilde{C}_{ha} = 0$, or with the bulk curve τ^b . In order to clearly demonstrate the effect of the surface contributions on the shape of the curve $\tau = \tau(H/p_0)$, the dimensionless ratio ϱ ,

$$\varrho = \frac{p_0}{T_0} \frac{\partial T}{\partial y} / \left(\frac{\partial p}{\partial y} \right), \quad (54)$$

of the transverse temperature gradient and the transverse pressure gradient is plotted versus H/p_0 in Figs. 3 and 4 for N₂ and HD, respectively. For both gases, ϱ is positive and of order unity. The surface contributions influence the value of ϱ in a different way at low and high fields; e.g. for the bulk contributions alone, ϱ decreases at higher H/p_0 , whereas the inclusion of the surface contributions may result in an increase of ϱ .

With the parameters used here, the total transverse temperature gradient for $\chi = 90^\circ$ has the same

Fig. 1. Bulk (····) and surface (—) contributions to the total (—) reduced temperature gradient τ for N₂, calculated for $\chi = 90^\circ$ from Eqs. (47), (48), (50), (51) with the help of Table I, $\tilde{C}_a = 1.0$, $\tilde{C}_{ma} = -0.3$ and three $\tilde{C}_{ha}$ values. The effective thermal slip number $\sigma + \alpha = 0.189$ was taken from Ref. ¹⁶.

Fig. 2. Same as Fig. 1 for the gas HD with $\tilde{C}_a = 0.4$, $\tilde{C}_{ma} = 0.4$, $\sigma + \alpha = 0.2$.

Fig. 3. The ratio $\varrho = \frac{p_0}{T_0} \frac{\partial T}{\partial y} / \left(\frac{\partial p}{\partial y} \right)$ for N₂, calculated for $\chi = 90^\circ$ from Eqs. (46)–(49) with the parameters from Fig. 1; ··· is the bulk contribution, — is the total effect for three $\tilde{C}_{ha}$ values.

Fig. 4. Same as Fig. 3 for HD with parameters from Figure 2.

Fig. 1.

Fig. 3

Fig. 2

Fig. 4

Fig. 5. Direction of the transverse temperature gradient $e_y \frac{\partial T}{\partial y}$ for N_2 ($g_r < 0$), with the magnetic field $\mathbf{H} = e_z H$ ($\chi = 90^\circ$) and the field free pressure gradient $\nabla p^{(0)} = e_x \left(\frac{\partial p}{\partial x} \right)_0$.

direction as the bulk contribution, viz. that of

$$\text{sign}(g_r) \nabla p^{(0)} \times \mathbf{h}.$$

The relation between the three vectors $e_y \frac{\partial T}{\partial y}$, $\nabla p^{(0)}$ and $\mathbf{h}$ is shown in Fig. 5 for N_2 .

Finally, the angular dependence of the effect shall be discussed, i. e. the change in $\partial T / \partial y$ through the variation of the field direction $\mathbf{h}$ in the plane perpendicular to the field-free flow direction e_x . If the angle χ gets smaller than 90° , the absolute height of the effect decreases, and the maximum occurs at lower H/p_0 -values. This can be easily understood by inspection of Eq. (50): $|\tau|$ decreases with about $\sin \chi$, and the double frequencies $2\varphi_{pq}$ become important at $\chi < 90^\circ$. In Fig. 6, the change of

the whole $\tau(H/p_0)$ curve with the angle χ is shown for N_2 and four χ -values. For the same gas, Figs. 7 and 8 display the dependence on χ of the value $|\tau|_{\max}$ and the position $(H/p_0)_{\max}$ of the maximum of $|\tau|$, both normalized with the $\chi = 90^\circ$ values. The $\chi = 90^\circ$ values of $|\tau|_{\max}$ and $(H/p_0)_{\max}$ for N_2 are listed in Table II for $\tilde{C}_a = 1.0$, two "reasonable" values¹² of $\tilde{C}_{ma}$ and four $\tilde{C}_{ha}$ values. The maximum value of $|\tau|$ increases with increasing $\tilde{C}_{ma}$ and decreasing $\tilde{C}_{ha}$ and shifts to higher H/p_0 values. From $|\tau|_{\max}$, $|\partial T / \partial y|_{\max}$ is easily obtained by observing that, with $T_0 = 300$ K, $p_0 = 1$ torr and $(\partial p / \partial x)_0 = 0.1$ torr/cm, a value of $\tau = 10^{-3}$ corresponds to $\partial T / \partial y = 30$ mK/cm.

Table II. Maximum position $(H/p_0)_{\max}$ and value $|\tau|_{\max}$ of the reduced temperature gradient τ for N_2 at $\chi = 90^\circ$, calculated from Eqs. (47), (48), (50) with the parameters listed in Table I, with $\tilde{C}_a = 1.0$ and two $\tilde{C}_{ma}$ -values¹², for four different $\tilde{C}_{ha}$. The effective thermal slip number $\sigma + \alpha = 0.189$ was taken from thermal transpiration data¹⁶.

$\tilde{C}_{ha}$	$\tilde{C}_{ma} = -0.3$		$\tilde{C}_{ma} = -0.8$	
	$(H/p_0)_{\max}$ kOe/torr	$10^3 \tau _{\max}$	$(H/p_0)_{\max}$ kOe/torr	$10^3 \tau _{\max}$
0	3.77	1.29	3.56	1.13
-1.0	3.89	1.64	3.71	1.45
-3.0	4.05	2.36	3.88	2.09
-5.0	4.13	3.07	3.99	2.74

Fig. 6. Reduced transverse temperature gradient τ for N_2 at four angles χ calculated from Eqs. (47), (48), (50) with $\tilde{C}_a = 1.0$, $\tilde{C}_{ma} = -0.3$, $\tilde{C}_{ha} = -3.0$, $\sigma + \alpha = 0.189$ and the parameters from Table I.

Fig. 7. Height of the maximum of $\left| \frac{\partial T}{\partial y} \right|$, normalized to the $\chi=90^\circ$ value, as a function of the angle χ for N_2 , $R_h = \left| \frac{\partial T}{\partial y} \right|_{\max}^{\chi} / \left| \frac{\partial T}{\partial y} \right|_{\max}^{90}$; the function $\sin \chi$ is indicated by $\cdots$. The parameters are $\tilde{C}_a=1.0$, $\tilde{C}_{ma}=-0.3$, $\sigma+\alpha=0.189$, $\tilde{C}_{ha}=0$ and $\tilde{C}_{ha}=-3.0$.

The existence of the transverse temperature gradient has been detected by Eggermont et al.⁹. A detailed experimental investigation of the effect is still going on. The comparison of these results with

Fig. 8. Position of the maximum of $\left| \frac{\partial T}{\partial y} \right|$, normalized to the $\chi=90^\circ$ value, as a function of the angle χ for N_2 , $R_p = (H/p_0)_{\max}^{\chi} / (H/p_0)_{\max}^{90}$. The parameters are the same as in Figure 7.

the presented theory will give information on the hitherto unknown surface parameter $\tilde{C}_{ha}$.

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